

1 **Preventing catastrophic filter divergence using adaptive additive inflation**
2 **for baroclinic turbulence**

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ABSTRACT

9 Ensemble based filtering or data assimilation methods have proved to be
10 indispensable tools in atmosphere and ocean science as they allow computa-
11 tionally cheap, low dimensional ensemble state approximation for extremely
12 high dimensional turbulent dynamical systems. For sparse, accurate and in-
13 frequent observations, which are typical in data assimilation of geophysical
14 systems, ensemble filtering methods can suffer from catastrophic filter diver-
15 gence which frequently drives the filter predictions to machine infinity. A
16 two-layer quasi-geostrophic equation which is a classical idealized model for
17 geophysical turbulence is used to demonstrate catastrophic filter divergence.
18 The mathematical theory of adaptive covariance inflation by *Tong et al.* and
19 covariance localization are investigated to stabilize the ensemble methods and
20 prevent catastrophic filter divergence. Two forecast models, a coarse-grained
21 ocean code, which ignores the small-scale parameterization, and stochastic
22 superparameterization (SP), which is a seamless multi-scale method devel-
23 oped for large-scale models without scale-gap between the resolved and unre-
24 solved scales, are applied to generate large-scale forecasts with a coarse spa-
25 tial resolution 48×48 compared to the full resolution 256×256 . The methods
26 are tested in various dynamical regimes in ocean with jets and vorticities, and
27 catastrophic filter divergence is documented for the standard filter without in-
28 flation. Using the two forecast models, various kinds of covariance inflation
29 with or without localization are compared. It shows that proper adaptive ad-
30 ditive inflation can effectively stabilize the ensemble methods without catas-
31 trophic filter divergence in all regimes. Furthermore, stochastic SP achieves
32 accurate filtering skill with localization while the ocean code performs poorly
33 even with localization.

34 **1. Introduction**

35 Ensemble based filters or data assimilation methods, including the ensemble Kalman filter
36 (EnKF; Evensen 2003) and ensemble square root filters, the ensemble transform Kalman filter
37 (ETKF; Bishop et al. 2001) and the ensemble adjustment Kalman filter (EAKF; Anderson 2001),
38 provide accurate statistical estimation of a geophysical system combining a forecast model and
39 observations. These methods quantify the uncertainty of the system using an ensemble which sam-
40 ples the information of the system. For geophysical systems which are complex high-dimensional
41 and thus require enormously huge computational costs for long time integration, the ensemble
42 based methods are indispensable tools for data assimilation as the methods allow computation-
43 ally cheap and low dimensional state approximation. Due to the simplicity and efficiency of the
44 ensemble based filters, these methods are widely applied to various fields of geophysical science
45 such as numerical weather prediction (Kalnay 2003).

46 Despite their successful applications in geophysical applications, ensemble based filters suffer
47 from small ensemble size due to the high dimensionality and expensive computational costs (fre-
48 quently referred as “curse of ensemble numbers” (Majda and Harlim 2012)) which can lead to
49 filter divergence. Sampling errors due to insufficient ensemble size and imperfect model errors
50 often yield underestimation of the uncertainty in the forecast and thus filters trust the forecast with
51 larger confidence than the information given by observations. Inaccurate uncertainty quantifica-
52 tion in the forecast fails to track the true signal and thus filter performance degrades, which is
53 called filter divergence (Majda and Harlim 2012). Also insufficient ensemble size can lead to spu-
54 rious overestimation of cross correlations between otherwise uncorrelated variables (Hamill et al.
55 2001; Whitaker et al. 2009; Sakov and Oke 2008) which also affect filter performance. Covari-
56 ance inflation, which inflates the prior covariance and pulls the filter back toward observations,
57 is one among various methods to remedy the filter divergence (Anderson 2001). For the over-
58 estimation of cross correlations between uncorrelated variables, localization which multiplies the
59 covariances between prior state variables and observation variables by a correlation function with

60 local support is a powerful method to correct the overestimated cross correlations (Houtekamer
61 and Mitchell 2001).

62 Catastrophic filter divergence (Harlim and Majda 2010; Gottwald and Majda 2013) is another
63 important issue hindering the applications of the ensemble based methods to high dimensional sys-
64 tems especially in the case with sparse and infrequent observations and small observation errors.
65 Catastrophic filter divergence drives the filter predictions to machine infinity although the under-
66 lying system remains in a bounded set. In data assimilation of geophysical systems in the ocean,
67 observations are often sparse and infrequent. In observations of ocean dynamics such as sea sur-
68 face temperature, observations become accurate using various techniques such as tropical moored
69 buoys, ocean reference status, and surface drifting buoys. But observations are still inadequate and
70 sparse to sample over the vast surface and the interior of the ocean.

71 It is shown rigorously in Kelly et al. 2015 that catastrophic filter divergence is not caused by
72 numerical instability, instead the analysis step of filters generates catastrophic filter divergence.
73 Although covariance inflation and localization stabilize filters and improve accuracy, they cannot
74 avoid catastrophic filter divergence. In Harlim and Majda 2010, it is demonstrated that ensemble
75 based methods with constant covariance inflation still suffer from catastrophic filter divergence. In
76 this study we also see that covariance localization decreases the occurrence of catastrophic filter
77 divergence but does not prevent catastrophic divergence.

78 To avoid catastrophic filter divergence, a judicious model error using linear stochastic models
79 was studied in Harlim and Majda 2010 with skillful results in some parameter regimes. Recently
80 a simple remedy of catastrophic filter divergence without using linear stochastic models has been
81 proposed through rigorous mathematical arguments and tested for the Lorenz-96 model in Tong
82 et al. 2016. The approach in Tong et al. 2016 adaptively inflates covariance with minimal addi-
83 tional costs according to the distribution of the ensemble. The strength of inflation is determined
84 by two statistics of the ensemble, 1) ensemble innovation which measures how far predicted ob-
85 servations are from actual observations and 2) cross covariance between observed and unobserved

86 variables (see (12) and (13) in Section 3 respectively). If the filter is malfunctioning based on these
87 two statistics, inflation is triggered and becomes larger when filters stray further into malfunction.

88 In this study we demonstrate catastrophic filter divergence of ensemble based filters in the two-
89 layer quasi-geostrophic equations, which are classical idealized models for geophysical turbulence
90 (Salmon 1998). The adaptive inflation method is then proposed for this two-layer system to avoid
91 catastrophic filter divergence. Both a coarse-grained ocean code, which ignores the subgrid scale
92 parameterization, and stochastic superparameterization (Grooms et al. 2015b), which is a seamless
93 multi-scale method developed for large-scale models without scale-gap between the resolved and
94 unresolved scales, are applied to generate forecasts with a coarse spatial resolution 48×48 for each
95 layer compared to the full resolution 256×256 which generates true signals. We test ensemble
96 methods for various dynamical regimes in the ocean corresponding to idealized low, mid and
97 high latitude states and document that catastrophic filter divergence occurs for ensemble based
98 methods even with localization unless adaptive inflation is applied. Ensemble filtering for the
99 two-layer quasi-geostrophic equations using these forecast models, the ocean code and stochastic
100 superparameterization, has already been studied in Grooms et al. 2015a to investigate the effect
101 of constant inflation on accounting for model errors without catastrophic filter divergence. In this
102 study we test a very sparse observation network which observes only 4×4 points of the upper
103 layer stream function with a small observation error variance corresponding to 1% of the total
104 variance of the stream function to represent the typical realistic scenario with sparse high quality
105 data, which leads to catastrophic filter divergence.

106 Using both the ocean code and stochastic superparameterization, various kinds of covariance
107 inflation with or without localization are compared. We verify that proper adaptive covariance
108 inflation can effectively stabilize the ensemble based filters uniformly without catastrophic filter
109 divergence in all test regimes. Furthermore, stochastic superparameterization achieves accurate
110 filtering skill with localization while the ocean code performs poorly even with localization.

111 The structure of this paper is as follows. In Section 2 we briefly review an ensemble method,
112 the Ensemble Adjustment Kalman Filter (Anderson 2001) with covariance inflation and local-

113 ization. The adaptive inflation method to prevent catastrophic filter divergence is described in
 114 Section 3 including how to choose parameters of the adaptive method. In Section 4 the two-
 115 layer quasi-geostrophic equation with baroclinic instability is described and two coarse-grained
 116 forecast models, the ocean code and stochastic superparameterization, are explained. Numerical
 117 experiments with various inflation strategies with or without localization are reported in Section 5
 118 along with stabilized and improved filtering results using the adaptive inflation method. In Section
 119 6 we conclude this paper with discussion.

120 2. Ensemble Filtering

121 In this section we briefly describe the Ensemble Adjustment Kalman Filter (EAKF; Anderson
 122 2001) which in our experience is a more stable and accurate scheme than other popular ensem-
 123 ble based methods (Majda and Harlim 2012). We assume that the true signal is generated by a
 124 nonlinear mapping $\psi_n : \mathbb{R}^d \rightarrow \mathbb{R}^d$

$$u_n = \psi_n(u_{n-1}) \quad (1)$$

125 where $u_n \in \mathbb{R}^d$ is a state vector at the n -th observation time. We consider a linear observation of
 126 u_n by an observation operator $H : \mathbb{R}^d \rightarrow \mathbb{R}^q$ with a rank q

$$z_n = Hu_n + \xi_n \quad (2)$$

127 where ξ_n is a mean zero Gaussian noise with a variance σ independent in different times and space
 128 grid points. For an easy exposition of the adaptive inflation in Section 3 we use a decomposition
 129 of the state variable u_n into observed and unobserved variables $x_n \in \mathbb{R}^q$ and $y_n \in \mathbb{R}^{d-q}$ respectively
 130 so that $x_n = Hu_n$ and $y_n \in \text{Ker}(H)$.

131 As other ensemble based filters, EAKF uses ensemble members $\{v_n^{(k)}\}_{k=1}^K$ to represent statis-
 132 tical properties of the state but uses only the first and second order moments (that is, mean and
 133 covariance) to update each ensemble member. First, EAKF generates prior predictions by solving

134 a forecast model for each ensemble member

$$\tilde{v}_n^{(k)} = \tilde{\psi}_n(v_{n-1}^{(k)}), \quad k = 1, 2, \dots, K \quad (3)$$

135 where $\tilde{\psi}_n$ is an approximate forecast model to the true dynamics ψ_n . From the forecast ensemble
 136 $\{\tilde{v}_n^{(k)}\}_{k=1}^K$, the prior mean \bar{v}_n^f and covariance C_n^f are given by

$$\bar{v}_n^f = \frac{1}{K} \sum_{k=1}^K \tilde{v}_n^{(k)} \quad (4)$$

137 and

$$C_n^f = \frac{1}{K-1} \sum_{k=1}^K \left(\tilde{v}_n^{(k)} - \bar{v}_n^f \right) \otimes \left(\tilde{v}_n^{(k)} - \bar{v}_n^f \right) \quad (5)$$

138 respectively. With these prior mean and covariance, the standard Kalman formula using observa-
 139 tion $z_n \in \mathbb{R}^q$ gives the following posterior mean and covariance

$$\bar{v}_n^a = \bar{v}_n^f - C_n H^T (I + H^T C_n H)^{-1} (H \bar{v}_n^f - z_n). \quad (6)$$

140 and

$$C_n^a = C_n^f - C_n^f H^T (I + H^T C_n^f H)^{-1} H C_n^f \quad (7)$$

141 respectively. For a ensemble perturbation matrix $V \in \mathbb{R}^{d \times K}$ whose k -th column is given by the
 142 ensemble perturbation $\delta v_n^{(k)} = \tilde{v}_n^{(k)} - \bar{v}_n^f$, EAKF finds an adjustment matrix $A_n \in \mathbb{R}^{K \times K}$ so that the
 143 adjusted ensemble satisfies the posterior covariance (7)

$$\frac{1}{K-1} V_n A_n \otimes V_n A_n = C_n - C_n H^T (I + H^T C_n H)^{-1} H C_n. \quad (8)$$

144 Once the adjustment matrix A_n is calculated, the posterior ensemble is obtained by adding the
 145 adjusted perturbation to the posterior mean. That is $v_n^{(k)} = \bar{v}_n^a + s_n^{(k)}$ where $s_n^{(k)}$ is the k -th column
 146 of the ensemble perturbation matrix $V_n A_n$.

147 Covariance inflation overcomes some problems caused by sampling errors due to insufficient
 148 ensemble numbers or an imperfect model and requires only a minimal additional cost to the origi-
 149 nal EAKF. The covariance inflation introduces more uncertainty in the prior covariance so that the
 150 filter has more weight on the information given by observations. That is, for a constant λ_n which
 151 determines the strength of inflation, covariance inflation inflates the prior covariance

$$C_n^f \leftarrow (I + \lambda_n)C_n^f. \quad (9)$$

152 for multiplicative inflation or

$$C_n^f \leftarrow C_n^f + \lambda_n I \quad (10)$$

153 for additive inflation. Then the ensemble is modified to satisfy the inflated prior covariance by
 154 spreading the ensemble for the multiplicative inflation and by adding additional noise for the addi-
 155 tive inflation. Although covariance inflation improves filter skill in many applications it is reported
 156 that constant inflation does not prevent catastrophic filter divergence with sparse and accurate ob-
 157 servation networks (Harlim and Majda 2010).

158 3. Adaptive Additive Inflation

159 A simple remedy in Tong et al. 2016 to stabilize ensemble based filters by preventing catas-
 160 trophic filter divergence is to adaptively trigger the inflation and change the strength λ_n . Although
 161 the adaptive inflation method of Tong et al. 2016 works both for the multiplicative inflation (9)
 162 and additive inflation (10), we focus on the simpler additive inflation in this study. The inflation
 163 strength λ_n of (10) is determined by two statistics of the ensemble

$$\lambda_n = c_a \Theta_n (1 + \Xi_n) \mathbb{1}_{\{\Theta_n > M_1 \text{ or } \Xi_n > M_2\}} \quad (11)$$

164 where c_a is a tunable positive constant, Θ_n is a measure related to the innovation process $H\tilde{v}_n^{(k)} - z_n$
 165 in a standard Kalman filter

$$\Theta_n := \frac{1}{K} \sum_{k=1}^K \|H\tilde{v}_n^{(k)} - z_n\|^2, \quad (12)$$

166 Ξ_n is the l_2 norm of the cross covariance between the observed and unobserved variables

$$\Xi_n = \left\| \frac{1}{K-1} \sum_{k=1}^K \left(\tilde{x}_n^{(k)} - \bar{x}_n \right) \otimes \left(\tilde{y}_n^{(k)} - \bar{y}_n \right) \right\|, \quad \tilde{v}_n^{(k)} = (\tilde{x}_n^{(k)}, \tilde{y}_n^{(k)}), \tilde{x}_n^{(k)} = H\tilde{v}_n^{(k)} \quad (13)$$

167 and M_1 and M_2 are fixed positive thresholds to decide whether the filter is performing well or not.
 168 The first statistical information Θ_n measures the accuracy of the prediction, that is, how far the
 169 predicted observations are from actual observations. The second statistical information Ξ_n is an
 170 important factor because large cross covariance can magnify a small error in the observed com-
 171 ponent and impose it on the unobserved variables. Hence the adaptive inflation can be regarded
 172 as a control of these two statistics to prevent catastrophic filter divergence. Note that these two
 173 factors are in fact derived from a rigorous mathematical argument for nonlinear stability of finite
 174 ensemble filters which can be found in Tong et al. 2016.

175 In contrast to the conventional covariance inflation which modifies the prior ensemble to satisfy
 176 the inflated covariance, the EAKF with adaptive additive inflation does not modify the prior en-
 177 semble to inflate covariance; the additive inflation can make the rank of the posterior covariance
 178 larger than or equal to d while its rank cannot exceed $K - 1$ where K is the ensemble size. Thus,
 179 in adaptive additive inflation, we use the inflated prior covariance (10) to calculate the posterior
 180 mean while the posterior covariance does not change. That is, instead of (6), the posterior mean is
 181 defined as

$$\bar{v}_n^a = \bar{v}_n^f - \tilde{C}_n^f H^T (I + H^T \tilde{C}_n^f H)^{-1} (H\bar{v}_n^f - z_n). \quad (14)$$

182 using

$$\tilde{C}_n^f = C_n^f + \lambda_n I \quad (15)$$

183 where the posterior covariance is the same as (7), that is, no inflated prior covariance.

184 The two thresholds M_1 and M_2 of (11) are important factors as they differentiate poor forecasts
 185 from properly working forecasts. Using an elementary benchmark of accuracy which should be
 186 surpassed by filters, we use the following aggressive thresholding (Tong et al. 2016). The thresh-
 187 olds are given by

$$M_1 = \|H\|^2 \text{Error}_{bench} + 2q\sigma \quad (16)$$

188 and

$$M_2 = \frac{K}{2K-2} \text{Error}_{bench} \quad (17)$$

189 where the benchmark for accuracy Error_{bench} is the mean-square error of an estimator using an
 190 invariant probability measure of the model

$$\text{Error}_{bench} := \mathbb{E}(u_n - \mathbb{E}(u_n|z_n))^2. \quad (18)$$

191 As the invariant measure of the model is not available, aggressive thresholding uses a Gaussian ap-
 192 proximation to the invariant measure using climatological properties, mean and covariance. Then
 193 the conditional distribution given observation z_n is a Gaussian measure and can be computed ex-
 194 actly which gives the following formula

$$\text{Error}_{bench} = \text{tr} \left(\text{cov}(u_n) - \text{cov}(u_n)H^T(I + H\text{cov}(u_n)H^T)^{-1}H\text{cov}(u_n) \right). \quad (19)$$

195 **4. Model equations and forecast models**

196 In atmosphere and ocean science, quasi-geostrophic equations are widely used as classical ide-
 197 alized models of geophysical turbulence (Salmon 1998). In this study we use a two-layer quasi-
 198 geostrophic equation as the model equation to observe catastrophic filter divergence in high di-
 199 mensional data assimilation and test the adaptive additive inflation to prevent catastrophic filter
 200 divergence. The system is maintained by baroclinic instability imposed by vertical shear flows

201 and shows interesting features in geophysical turbulence such as inverse cascade of energy and
 202 zonal jets. After describing the model equation in Section a, two coarse-grained forecast mod-
 203 els, an ocean code which ignores the subgrid scales and another forecast method with stochastic
 204 parameterization of the subgrid scales, are explained in Section b.

205 *a. Two-layer quasi-geostrophic equations*

206 Our model equation to generate high dimensional geophysical turbulence is the following two-
 207 layer quasi-geostrophic equation in a doubly periodic domain used in Grooms and Majda 2014;
 208 Majda and Grooms 2014; Grooms et al. 2015a; Lee et al. 2016 to generate baroclinic turbulence

$$\begin{aligned}
 \partial_t q_1 &= -\mathbf{v}_1 \cdot \nabla q_1 - \partial_x q_1 - (k_\beta^2 + k_d^2)v_1 - \nu \Delta^4 q_1, \\
 \partial_t q_2 &= -\mathbf{v}_2 \cdot \nabla q_2 + \partial_x q_2 - (k_\beta^2 - k_d^2)v_2 - r \Delta \psi_2 - \nu \Delta^4 q_2, \\
 q_1 &= \Delta \psi_1 + \frac{k_d^2}{2}(\psi_2 - \psi_1), \\
 q_2 &= \Delta \psi_2 - \frac{k_d^2}{2}(\psi_2 - \psi_1).
 \end{aligned} \tag{20}$$

209 Here q_j is the potential vorticity in the upper ($j = 1$) and lower ($j = 2$) layers, k_d is the defor-
 210 mation wavenumber, r is a linear Ekman drag coefficient at the bottom layer of the flows, k_β is
 211 an nondimensional constant resulting from the variation of the vertical projection of Coriolis fre-
 212 quency with latitude and the velocity field $\mathbf{v}_j = (u_j, v_j) = (-\partial_y \psi_j, \partial_x \psi_j)$ for the stream function
 213 ψ_j . To stabilize the equation by absorbing a downscale cascade of enstrophy at the smallest scales
 214 while leaving other scales nearly inviscid for interesting dynamics at large-scales, we use a hyper-
 215 dissipation $\Delta^4 q_j$ with a hyperviscosity ν , which is commonly used in turbulence simulations. To
 216 maintain nontrivial dynamics of (20) by baroclinic instability, a large-scale zonal vertical shear is
 217 applied with equal and opposite unit velocities which are related to the terms $(-1)^j(\partial_x q_j + k_d^2 v_j)$
 218 in (20).

219 Following the experiments in Grooms et al. 2015a and Lee et al. 2016, we test three different
220 regimes corresponding to low, mid and high latitude ocean models by changing the β -plane effect
221 k_β and the bottom drag r (see Table 1 for the parameter values of the three test regimes). While
222 the deformation wavenumber k_d is fixed at 25, we use a fine resolution of 256×256 grid points
223 for each layer to generate true signals in our data assimilation experiments. The hyperviscosity ν
224 is set to 1.28×10^{-15} and we use a pseudo-spectral space discretization while the time integration
225 uses a fourth order semi-implicit Runge-Kutta method by incorporating an exponential integration
226 for the linear stiff dissipation term. Time step is fixed at 2×10^{-5} for all test regimes.

227 In the high latitude case (or the f -plane case), the quasi-geostrophic equation is dominated by
228 spatially homogeneous and isotropic flows (see Figure 1 for snapshots of the upper and lower
229 layer stream function). In the mid and low latitude cases which have the β -plane effect, the flows
230 organize into inhomogeneous and anisotropic structure such as zonal jets.

231 *b. Forecast models with and without stochastic parameterization*

232 As a forecast model in data assimilation of the true signal given by (20), we consider two fore-
233 cast models on a low resolution 48×48 grid points, 1) an ocean code which uses only a coarse
234 grid without parameterizing the small scales and 2) stochastic superparameterization which pa-
235 rameterizes the effect of the small scales by modeling the small scales as randomly oriented plane
236 waves (Majda and Grooms 2014; Grooms et al. 2015b). Note that these two forecast models are
237 imperfect models as they approximate the true signal on a low resolution grid. Thus in data as-
238 simulation using ensemble based methods, there is an error from the imperfect model in addition
239 to the sampling error due to a small ensemble size.

240 The first forecast model, which we call the ocean code, solves the following approximation to
241 (20) which replaces the hyper dissipation by a biharmonic dissipation of relative vorticity $\omega_j =$

242 $\Delta\psi_j$

$$\begin{aligned}\partial_t q_1 &= -\mathbf{v}_1 \cdot \nabla q_1 - \partial_x q_1 - (k_\beta^2 + k_d^2)v_1 - v_2 \Delta^2 \omega_1, \\ \partial_t q_2 &= -\mathbf{v}_2 \cdot \nabla q_2 + \partial_x q_2 - (k_\beta^2 - k_d^2)v_2 - r\Delta\psi_2 - v_2 \Delta^2 \omega_2.\end{aligned}\tag{21}$$

243 This replacement is to mimic the biharmonic dissipation commonly used in eddy-permitting ocean
244 models (Griffies and Hallberg 2000). By analogy with ocean models and some atmospheric mod-
245 els, the ocean code also uses the second order energy- and enstrophy-conserving Arakawa finite
246 differencing (Arakawa 1966) for the nonlinear advection terms $\mathbf{v}_j \cdot \nabla q_j, j = 1, 2$. For time integra-
247 tion, we use a second order Runge-Kutta integration with the same exponential integrator for the
248 linear stiff term and a time step fixed at 5×10^{-4} .

249 We consider another forecast model called stochastic superparameterization which uses stochas-
250 tic parameterization of the subgrid scales using randomly oriented plane waves for the subgrid
251 scales. The subgrid scales are generally not zero and influence the evolution of the resolved scales.
252 Especially in quasi-geostrophic turbulence which includes regimes with a net transfer of kinetic
253 energy from small to large scales (Charney 1971), it is important to accurately model the effects
254 of the under-resolved eddies to obtain accurate properties of the system such as energy spectrum.
255 Stochastic superparameterization is developed as a multiscale model for turbulence without scale-
256 gap between the resolved and unresolved scales (Grooms and Majda 2014; Majda and Grooms
257 2014). Among various versions of stochastic superparameterization, we use the most recent ver-
258 sion developed in Grooms et al. 2015b to deal with arbitrary boundary conditions using finite
259 difference numerics for the large scales.

260 The stochastic superparameterization forecast model solves (21) using the same second order
261 finite differencing for the nonlinear term but with additional terms $SGS_j, j = 1, 2$ obtained from

262 stochastic subgrid scale parameterization

$$\begin{aligned}
\partial_t q_1 &= -\mathbf{v}_1 \cdot \nabla q_1 - \partial_x q_1 - (k_\beta^2 + k_d^2) v_1 - \nu_2 \Delta^2 \omega_1 + SGS_1, \\
\partial_t q_2 &= -\mathbf{v}_2 \cdot \nabla q_2 + \partial_x q_2 - (k_\beta^2 - k_d^2) v_2 - r \Delta \psi_2 - \nu_2 \Delta^2 \omega_2 + SGS_2.
\end{aligned} \tag{22}$$

263 The parameterization terms $SGS_j, j = 1, 2$ are computed by modeling the subgrid scale as ran-
264 domly oriented plane waves. Under this modeling of the subgrid scales, stochastic superparame-
265 terization replaces the nonlinear terms of the subgrid scale equation using additional damping and
266 white noise forcing which yields quasilinear equation conditional to the resolved scale variable.
267 We also use the method in Grooms et al. 2015b to impose temporal correlations in the parameteri-
268 zation by using a Wiener process model for the orientation of the plane waves. Because the subgrid
269 scales are solved in formally infinite domains, this approach has no scale-gap between the resolved
270 and subgrid scales. Also, the stochastic modeling of the subgrid scales generate the missing insta-
271 bility of the subgrid scales using deterministic parameterization of the subgrid scales. Note that
272 we use the same time integration as in the ocean code, thus the difference between the ocean code
273 and stochastic superparameterization comes from the parameterization terms $SGS_j, j = 1, 2$.

Figure 2 shows the time averaged kinetic energy (KE) spectra

$$KE = \frac{1}{2} \int |\nabla \psi_1|^2 + |\nabla \psi_2|^2$$

274 by the direct numerical method (black), stochastic superparameterization (blue) and the ocean
275 code (red) using biharmonic viscosity $\nu_4 = 1.0 \times 10^{-7}$ and $\nu_4 = 1.6 \times 10^{-4}$ obtained by tuning to
276 match the energy spectra. Although the ocean code has much weaker dissipation than stochastic
277 superparameterization, the ocean code has smaller energies than stochastic superparameterization
278 while stochastic superparameterization captures the correct large-scale kinetic energy spectra; the
279 small energy of the ocean code cannot be improved further by tuning the biharmonic viscosity
280 coefficient. This result implies that the ocean code could have filter divergence by inappropriately

281 capturing the uncertainty in the forecast due to small energy of the resolved scales. On the other
282 hand, it is shown that stochastic parameterization can act to reduce model error (Shutts 2005;
283 Frenkel et al. 2012) and it increases ensemble spread which yields an effect similar to covariance
284 inflation. In the next section, we will see that stochastic superparameterization requires smaller
285 covariance inflation than the ocean code as the ocean code has large model errors which cannot be
286 improved by covariance inflation.

287 **5. Catastrophic filter divergence and numerical experiments**

288 In this section we demonstrate catastrophic filter divergence for all three test regimes regard-
289 less of the two forecast models, the ocean code and stochastic superparameterization, with sparse
290 high quality observations which are infrequent in time. Catastrophic filter divergence is effectively
291 prevented using the adaptive additive inflation for both forecast methods. Stochastic superparam-
292 eterization achieves accurate filtering skill with localization while the ocean code fails to achieve
293 accurate skill even with localization.

294 *a. Filtering setup*

295 For EAKF, we use a sequential update of observations used in Anderson 2001 which avoids
296 explicit computation of the SVD in (8) by processing observations individually. The true signal is
297 given by a fine resolution solution of (20) using a resolution 256×256 for each layer and a fourth
298 order semi-implicit Runge-Kutta integration with a time step 2×10^{-5} . The two forecast models,
299 the ocean code and stochastic superparameterization, use the same coarse resolution 48×48 for
300 each layer and a second order semi-implicit Runge-Kutta with a time step 5×10^{-4} .

301 We observe only the upper layer stream function, analogous to observation of sea surface height,
302 on a sparse 4×4 uniform grid while the stream function in the lower layer is completely unob-
303 served. Observation error variances correspond to 1% of the stream function variance for each test
304 regime. Following the idea of Keating et al. 2012, the eddy turnover time $T_{eddy} = 2\pi Z^{-1/2}$ (where
305 Z is the time-averaged total enstrophy $q_1^2 + q_2^2$) is comparable to 0.006 for all test regimes and we

306 use infrequent observations with an observation interval 0.008. Note that using the time step of the
 307 forecast models, 5×10^{-4} , this observation interval requires 16 time integrations for each forecast
 308 step. The ensemble size is 17 and thus the prior covariance is not necessarily rank deficient. This
 309 number is small compared to the dimension of the forecast model which is general in real data
 310 assimilation.

311 For each filtering test, we run 1000 assimilation cycles and take the last 600 cycles to measure
 312 filter performance using the time averaged RMS error (RMSE)

$$\text{time averaged RMSE} := \frac{1}{600} \sum_{n=401}^{1000} \|\bar{v}_n - u_n\| \quad (23)$$

313 and pattern correlation (PC)

$$\text{time averaged PC} := \frac{1}{600} \sum_{n=401}^{1000} \frac{\langle \bar{v}_n, u_n \rangle}{\|\bar{v}_n\| \|u_n\|} \quad (24)$$

314 respectively where $\langle \cdot, \cdot \rangle$ is the l_2 -inner product.

315 For covariance inflation, we test several combinations of inflation methods - for the inflation
 316 strength λ_n in (10), no inflation (noI) $\lambda_n = 0$, constant inflation (CI) $\lambda_n = c_c$ for a constant c_c ,
 317 adaptive inflation (AI) λ_n by (11) and constant+adaptive inflation (CAI)

$$\lambda_n = c_c + c_a \Theta_n (1 + \Xi_n) \mathbb{1}_{\{\Theta_n > M_1 \text{ or } \Xi_n > M_2\}} \quad (25)$$

318 (see Table 2 for the tuned c_c and c_a used in this study). The thresholds for adaptive inflation are
 319 given by the aggressive thresholding (16) and (17) where the benchmark for accuracy, Error_{bench} is
 320 given by 10, 166 and 155 for the low, mid and high latitude cases respectively from the reference
 321 simulations. Along with these inflation methods, we also use covariance localization with the
 322 compactly supported fifth-order piecewise rational function from Gaspari and Cohn 1999. The
 323 localization radius (where influence of observation is zero) is set to 8 forecast grid points. The

324 distance between two adjacent observation points are 12 and thus the square region centered at
325 each observation point is marginally updated from the other observation points.

326 *b. Filter experiments - catastrophic filter divergence and stabilization*

327 If no inflation is applied, EAKF has catastrophic filter divergence for both forecast models.
328 Figure 3 shows a sequence of snapshots of the low latitude case upper layer stream function by the
329 ocean code without inflation and localization (observation points are marked with black circles).
330 At the 570th cycle, the filter still works capturing the meridional structure of the low latitude case
331 but as more cycles go on, instability develops at unobserved grid points which eventually diverges
332 to machine infinity after the 600th cycle. The first row of Figure 4 shows time series of the RMS
333 errors by each forecast method when they suffer from catastrophic filter divergence. The RMS
334 errors increase gradually but they eventually diverge to machine infinity. The two forecast models
335 run slightly longer with localization but localization fails to prevent catastrophic filter divergence.
336 The second row of Figure 4 shows time series of RMS errors with the constant+adaptive inflation
337 where the cycles at which adaptive inflation is triggered is marked with dots. In the ocean code
338 case with no localization, the adaptive inflation is triggered at the beginning and stops although the
339 filter still degrades. Inflation is triggered again when the filter fails to capture the true signal. The
340 ocean code with localization triggers the adaptive inflation most of the time and obtains a stable
341 result but also fails to achieve accurate filtering skill. In the stochastic superparameterization case
342 with adaptive inflation and localization, adaptive inflation is triggered only 99 times out of 1000
343 cycles where most of the adaptive inflation is triggered at the beginning and infrequently triggered
344 later as the filter is performing well.

345 The occurrence percentage of catastrophic filter divergence out of 100 different runs is in Table
346 3. With no localization and inflation, the filter suffers from catastrophic filter divergence more than
347 75% for both the ocean code and stochastic superparameterization. The constant inflation stabi-
348 lizes the filter slightly but it does not prevent catastrophic filter divergence perfectly. The constant
349 inflation (CI) with no localization has a higher percentage of divergence than the no inflation case

350 for the stochastic superparameterization forecast model. Through stochastic parameterization of
351 subgrid scales, stochastic superparameterization has more variability than the ocean code and thus
352 additional constant inflation is not necessary.

353 Adaptive inflation (AI) with and without localization significantly decreases the number of oc-
354 currence of catastrophic filter divergence but the ocean code fails to prevent catastrophic filter
355 divergence entirely. For the constant+adaptive inflation (CAI), all methods are stable even without
356 localization. Note that for stochastic superparameterization, both AI and CAI work well pre-
357 venting catastrophic filter divergence while the ocean code fails to prevent the divergence in the
358 AI case. As we discussed before, stochastic superparameterization has enough ensemble spread
359 through stochastic parameterization of the subgrid scales and thus when adaptive inflation is al-
360 ready applied, constant inflation plays a marginal role in improving filter skill.

361 For the stabilized filters with the constant+adaptive inflation, we compare the filter performance
362 using the time averaged posterior RMS errors and pattern correlations (the performance difference
363 between the adaptive inflation (AI) and constant+adaptive inflation (CAI) is marginal when there is
364 no catastrophic filter divergence). In the low latitude case (shown in Table 4), both the ocean code
365 and the superparameterization methods fail to achieve accurate filtering skill without localization.
366 The RMS errors are larger than the standard deviation of the stream function and both forecast
367 methods do not capture the correlation with the true signal. When localization is combined with
368 adaptive inflation, it helps to increase filtering skill for both methods. The superparameterization
369 has significantly improved results; RMS error is smaller than 50% of the standard deviation of the
370 stream function and pattern correlation is larger than 90% for both layers. Although the lower layer
371 stream function is completely unobserved, the adaptive filter achieves accurate filter skill. The
372 ocean code result is improved using localization but it still suffers from standard filter divergence
373 with RMS error larger than the standard deviation of the stream function.

374 In the mid latitude case, the superparameterization still has skillful filtering skill and is superior
375 to the ocean code although the performance is slightly degraded compared to the low latitude case
376 as the mid latitude is more turbulent than the low latitude case. The RMS error by superparame-

377 terization with adaptive inflation and localization is about 30% smaller than the standard deviation
378 and pattern correlations are larger than 75% (see Table 5 for the mid latitude case RMS errors
379 and pattern correlations). On the other hand, the ocean code does not show any significant skill
380 even with adaptive inflation and localization. In the mid latitude case, the ocean code using adap-
381 tive inflation displays comparable results with and without localization, and both fail to achieve
382 meaningful filtering results. For the superparameterization, on the other hand, significantly im-
383 provement in filter skill can be achieved using localization (see the second row of Figure 5 the
384 time series of RMS errors with adaptive inflation). As the RMS errors are more fluctuating than
385 the low latitude case, the adaptive inflation is triggered most of the time for all combination of
386 inflation and localization.

387 The last test regime, high latitude case, is the most difficult test case as it is strongly turbulent
388 and dominated by homogeneous and isotropic vortical flows with no spatial structure. In this test
389 regime, stochastic superparameterization with constant+adaptive inflation and localization still
390 achieves a smaller RMS error and a larger pattern correlation than the ocean code though the
391 improvement by superparameterization is more marginal. The observed upper layer RMS error is
392 10% smaller than the standard deviation while the unobserved lower layer RMS error is only 5%
393 smaller than the standard deviation (Table 6).

394 **6. Conclusions**

395 Ensemble based filtering methods are indispensable tools in atmosphere and ocean science as
396 they provide computationally cheap and low dimensional ensemble state estimation for extremely
397 high dimensional turbulent systems. But these methods can suffer from catastrophic filter diver-
398 gence which drives the forecast predictions to machine infinity especially when the observation
399 is sparse, accurate and infrequent although the underlying true signal remains bounded. Using
400 an idealized model for the geophysical turbulence of the ocean, the two-layer quasi-geostrophic
401 equation with baroclinic instability, and a sparse observation network which is general in real ap-

402 plications, we were able to see catastrophic filter divergence of the ensemble adjustment Kalman
403 filter, which is one of the most stable and accurate ensemble methods.

404 The constant covariance inflation and localization, which are widely used methods to account
405 for the sampling errors due to insufficient ensemble size and model errors from imperfect forecast
406 models, stabilize the filter but fail to prevent the catastrophic filter divergence. Increasing the
407 observation size or ensemble number can help to prevent catastrophic filter divergence but this
408 approach is practically prohibitive and sometimes impossible as it requires enormous amount of
409 financial and computer resources to cover the vast surface of the ocean. Instead we followed
410 the adaptive inflation approach of Tong et al. 2016 to prevent catastrophic filter divergence. The
411 adaptive approach requires a minimal additional computational cost compared to the standard
412 ensemble based methods and uses only two low order statistics of the ensemble, the ensemble
413 innovation and cross covariance between observed and unobserved variables.

414 We tested the adaptive inflation using two forecast models, the ocean code without parameteri-
415 zation of the subgrid scales and stochastic superparameterization which parameterizes the subgrid
416 scales by modeling them as randomly oriented plane waves. Although both forecast models are
417 stabilized with the adaptive inflation, stochastic superparameterization displays filtering skill su-
418 perior to the ocean code. When the ensemble method is combined with localization and adaptive
419 inflation, stochastic superparameterization achieves RMS errors smaller than the climatological
420 error while the ocean code still suffers from the standard filter divergence with RMS errors com-
421 parable to the climatological error.

422 As we have shown in this study, covariance inflation is an important and useful technique in
423 ensemble based methods to improve filtering skill. There are another class of adaptive inflation
424 techniques such as Anderson 2007 and Ying and Zhang 2015. Although the adaptive inflation in
425 Tong et al. 2016 is based on rigorous mathematical arguments, it would be interesting to test other
426 adaptive inflation methods to avoid catastrophic filter divergence like the blended filter (Majda
427 et al. 2014; Qi and Majda 2015) that combines a particle filter in a low-dimensional subspace
428 and efficient Kalman filter in the orthogonal part. As it is investigated in Harlim and Majda 2010

429 through a linear stochastic model for the forecast, a judicious model error could be alternative to
430 prevent catastrophic filter divergence.

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499 **LIST OF TABLES**

500 **Table 1.** Parameters of (20) for three test regimes. Other parameters are fixed at $\nu =$
501 1.28×10^{-15} and $k_d = 25$ 26

502 **Table 2.** Constant and adaptive inflation parameters c_c and c_a for each test regime 27

503 **Table 3.** Occurrence percentage of catastrophic filter divergence out of 100 different
504 runs with and without localization. No inflation (noI), constant (CI), adaptive
505 (AI) and constant+adaptive (CAI) inflation methods. 28

506 **Table 4.** Low latitude case. Stream function estimation for both layers. Posterior RMS
507 errors and pattern correlations in parenthesis 29

508 **Table 5.** Mid latitude case. Stream function estimation for both layers. Posterior RMS
509 errors and pattern correlations in parenthesis 30

510 **Table 6.** High latitude case. Stream function estimation for both layers. Posterior RMS
511 errors and pattern correlations in parenthesis 31

TABLE 1: Parameters of (20) for three test regimes. Other parameters are fixed at $\nu = 1.28 \times 10^{-15}$ and $k_d = 25$

	k_β	r
Low	$k_d^2/2$	0.5
Mid	$k_d^2/4$	2
High	0	8

TABLE 2: Constant and adaptive inflation parameters c_c and c_a for each test regime

	Ocean		SP	
	c_c	c_a	c_c	c_a
Low	3×10^{-3}	5×10^{-4}	1×10^{-4}	5×10^{-4}
Mid	2×10^{-3}	4×10^{-4}	2×10^{-4}	1×10^{-5}
High	3×10^{-3}	1×10^{-4}	1×10^{-3}	1×10^{-5}

	Low		Mid		High	
no localization	Ocean	SP	Ocean	SP	Ocean	SP
noI	78%	84%	98%	97%	90%	85%
CI	63%	87%	80%	76%	45%	57%
AI	3%	0%	2%	0%	5%	0%
CAI	0%	0%	0%	0%	0%	0%
	Low		Mid		High	
with localization	Ocean	SP	Ocean	SP	Ocean	SP
noI	40%	24%	19%	38%	44%	64%
CI	15%	11%	9%	12%	22%	8%
AI	1%	0%	0%	0%	0%	0%
CAI	0%	0%	0%	0%	0%	0%

TABLE 3: Occurrence percentage of catastrophic filter divergence out of 100 different runs with and without localization. No inflation (noI), constant (CI), adaptive (AI) and constant+adaptive (CAI) inflation methods.

TABLE 4: Low latitude case. Stream function estimation for both layers. Posterior RMS errors and pattern correlations in parenthesis

no localization	Ocean, CAI	SP, CAI	Std of stream ftn
Upper layer	4.35 (0.05)	3.18 (0.02)	3.21
Lower layer	4.40 (0.01)	3.24 (0.04)	3.07
with localization	Ocean, CAI	SP, CAI	Std of stream ftn
Upper layer	3.76 (0.62)	1.37 (0.93)	3.21
Lower layer	3.90 (0.50)	1.41 (0.91)	3.07

TABLE 5: Mid latitude case. Stream function estimation for both layers. Posterior RMS errors and pattern correlations in parenthesis

no localization	Ocean, CAI	SP, CAI	Std of stream ftn
Upper layer	11.99 (0.26)	11.62 (0.30)	12.59
Lower layer	12.58 (0.23)	12.13 (0.29)	12.03
with localization	Ocean, CAI	SP, CAI	Std of stream ftn
Upper layer	11.24 (0.35)	7.73 (0.78)	12.59
Lower layer	12.26 (0.23)	8.14 (0.77)	12.03

TABLE 6: High latitude case. Stream function estimation for both layers. Posterior RMS errors and pattern correlations in parenthesis

no localization	Ocean, CAI	SP, CAI	Std of stream ftn
Upper layer	14.12 (0.25)	11.95 (0.22)	12.71
Lower layer	14.67 (0.24)	12.98 (0.18)	12.21
with localization	Ocean, CAI	SP, CAI	Std of stream ftn
Upper layer	13.01 (0.31)	11.38 (0.43)	12.71
Lower layer	13.21 (0.26)	11.53 (0.42)	12.21

512 **LIST OF FIGURES**

513 **Fig. 1.** Snapshots of stream functions $\psi_j, j = 1, 2$. Upper layer (top row) and lower layer (bottom
514 row). Low (first column), mid (second column) and high (third column) latitude cases. 33

515 **Fig. 2.** Time averaged total kinetic energy (KE) spectra by direct numerical reference (black),
516 stochastic superparameterization (blue) and ocean code (red) 34

517 **Fig. 3.** Low latitude case. Snapshots of posterior upper layer stream functions by Ocean code at
518 570th, 580th, 590th, and 600th cycles. Observation points are marked with circles. Catastrophic filter
519 divergence is invoked after the 600th cycle. 35

520 **Fig. 4.** Low latitude case. Time series of upper layer RMS error. The cycles at which inflation is
521 triggered are marked with filled circles. Standard deviation of the stream function in dash
522 line. 36

523 **Fig. 5.** Mid latitude case. Time series of upper layer RMS error. The cycles at which inflation is
524 triggered are marked with filled circles. Standard deviation of the stream function in dash
525 line. 37

526 **Fig. 6.** High latitude case. Time series of upper layer RMS error. The cycles at which inflation is
527 triggered are marked with filled circles. Standard deviation of the stream function in dash
528 line. 38

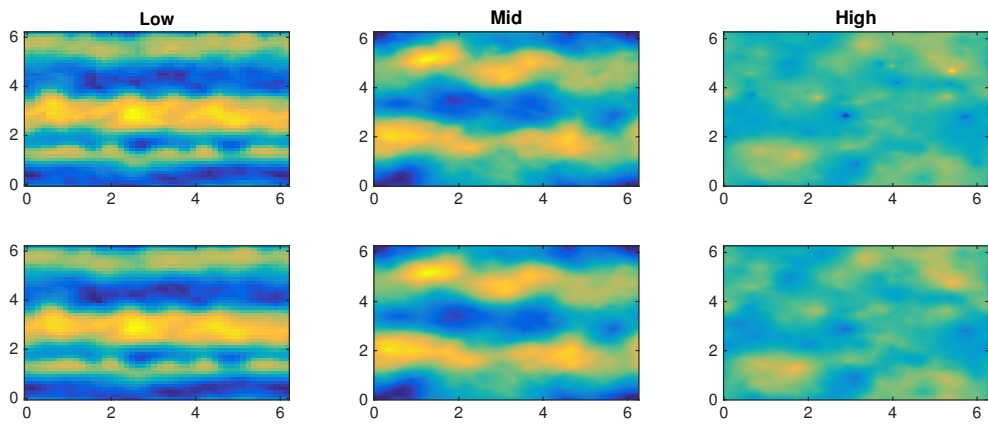


FIG. 1: Snapshots of stream functions $\psi_j, j = 1, 2$. Upper layer (top row) and lower layer (bottom row). Low (first column), mid (second column) and high (third column) latitude cases.

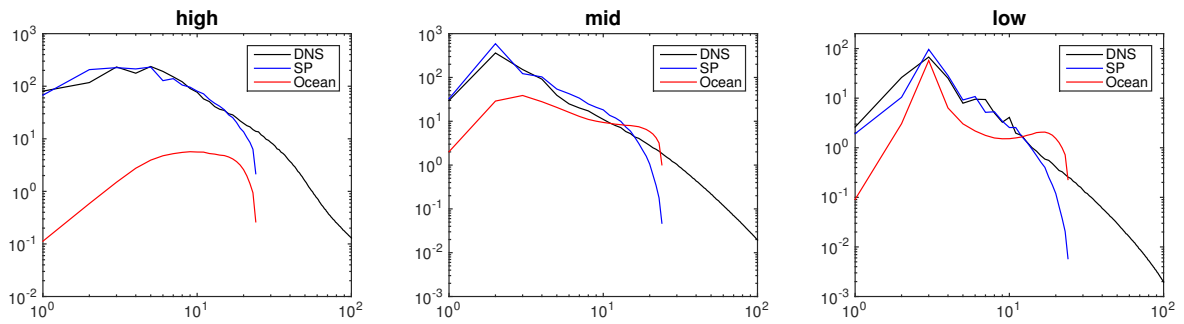


FIG. 2: Time averaged total kinetic energy (KE) spectra by direct numerical reference (black), stochastic superparameterization (blue) and ocean code (red)

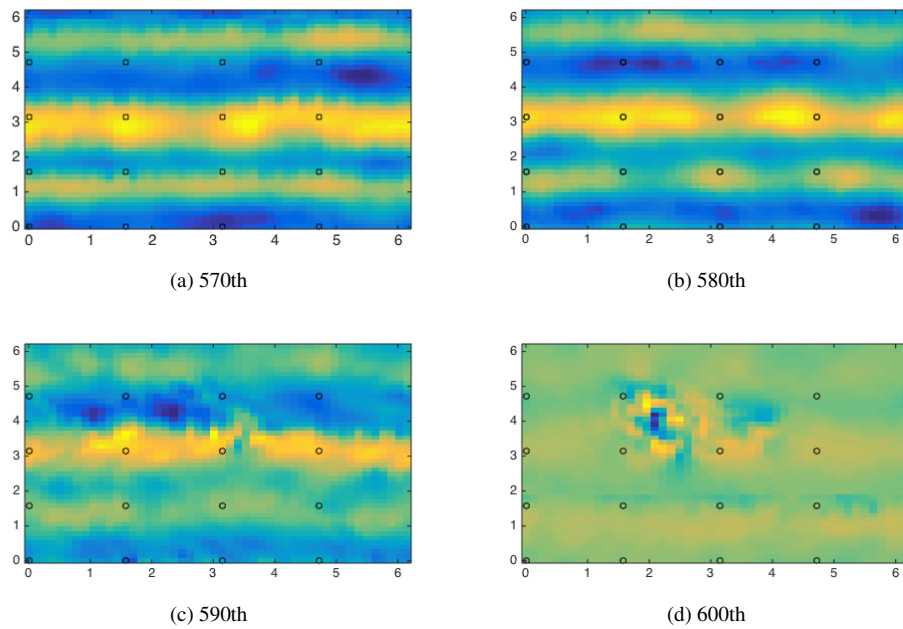


FIG. 3: Low latitude case. Snapshots of posterior upper layer stream functions by Ocean code at 570th, 580th, 590th, and 600th cycles. Observation points are marked with circles. Catastrophic filter divergence is invoked after the 600th cycle.

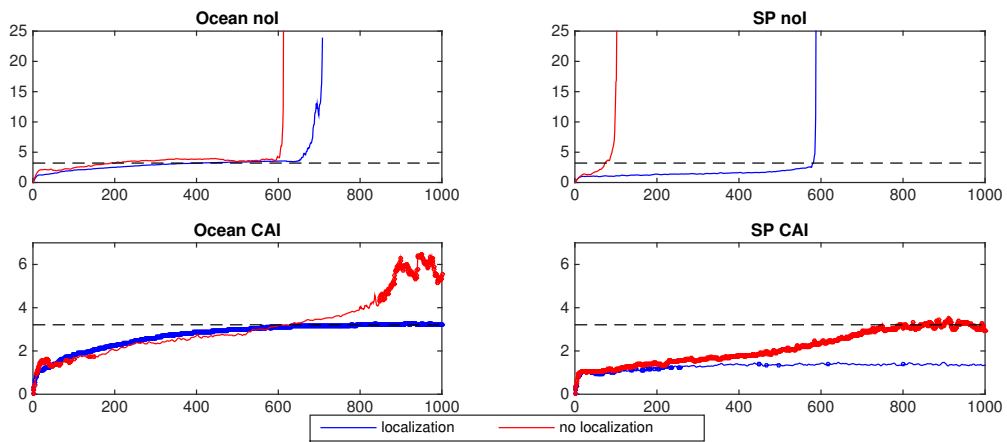


FIG. 4: Low latitude case. Time series of upper layer RMS error. The cycles at which inflation is triggered are marked with filled circles. Standard deviation of the stream function in dash line.

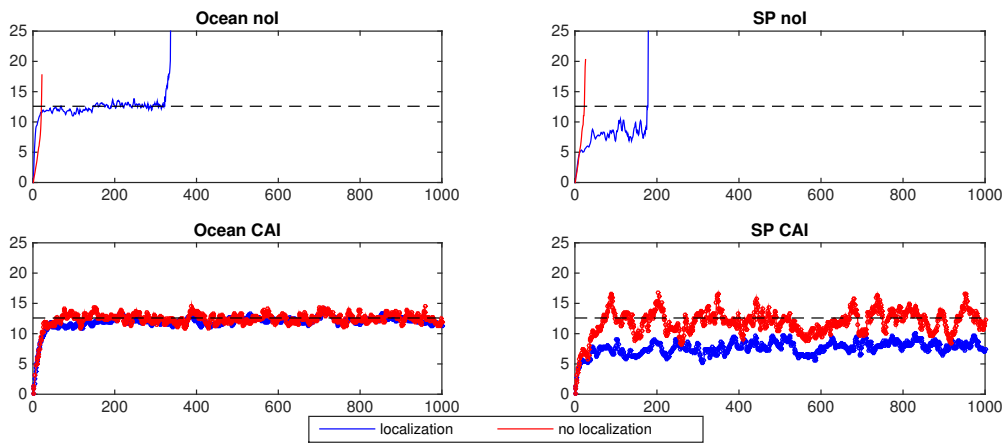


FIG. 5: Mid latitude case. Time series of upper layer RMS error. The cycles at which inflation is triggered are marked with filled circles. Standard deviation of the stream function in dash line.

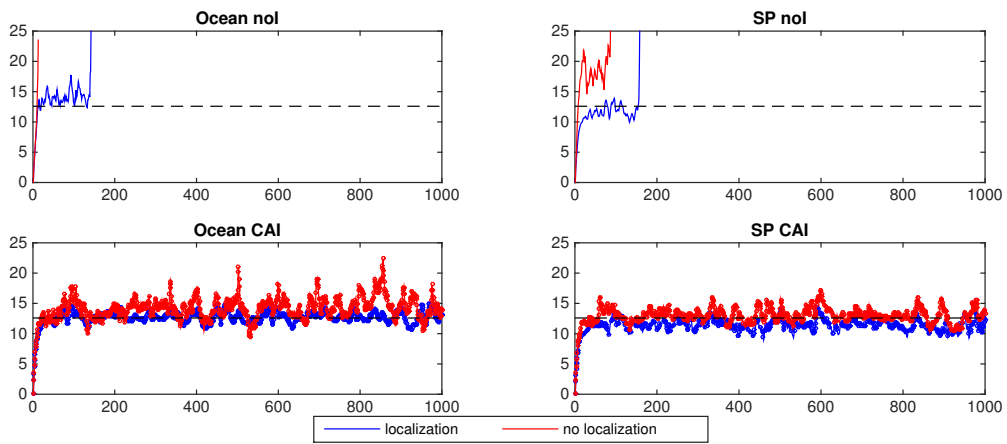


FIG. 6: High latitude case. Time series of upper layer RMS error. The cycles at which inflation is triggered are marked with filled circles. Standard deviation of the stream function in dash line.