

1 **Upscale Impact of Mesoscale Disturbances of Tropical Convection on**
2 **Synoptic-Scale Equatorial Waves in Two-Dimensional Flows**

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ABSTRACT

9 Superclusters of cloudiness on the synoptic scale are frequently organized
10 by convectively coupled equatorial waves (CCEWs). Within the large-scale
11 convective envelope, numerous mesoscale disturbances of tropical convec-
12 tion are typically found. It is a challenge for present-day numerical mod-
13 els to simulate such multi-scale structure of tropical convection. Also, the
14 upscale impact of mesoscale disturbances on the behavior of synoptic-scale
15 circulation is unclear. It is still not well understood how much of synoptic-
16 scale circulation is induced by upscale impact of mesoscale disturbances in-
17 stead of mean heating. Here a simple two-dimensional multi-scale model
18 for scale interactions across mesoscale and synoptic scale is used. A pre-
19 scribed two-scale heating drives synoptic-scale circulation through eastward-
20 moving mean heating and eddy transfer of momentum and temperature, where
21 the latter represents the upscale impact of mesoscale disturbances driven by
22 westward-moving mesoscale heating. This multi-scale model successfully re-
23 produces many key features of flow fields with a front-to-rear tilt, which are
24 compared with results from a cloud resolving model. In the scenario with
25 an elevated upright mean heating, the tilted vertical structure of synoptic-
26 scale circulation is still induced from the upscale impact of mesoscale distur-
27 bances. In the faster propagation scenario, the upscale impact becomes less
28 important due to competing effects of eddy transfer of momentum and tem-
29 perature, while the synoptic-scale circulation response to mean heating dom-
30 inates, in agreement with cloud resolving models. In the unrealistic scenario
31 with upward/westward tilted mesoscale heating, positive potential tempera-
32 ture anomalies are induced in the leading edge, which will suppress shallow
33 convection in a moist environment.

34 **1. Introduction**

35 Tropical convection is organized in a hierarchy of multiple spatial and temporal scales, ranging
36 from cumulus clouds over several kilometers to mesoscale circulation systems (MCSs) (Houze
37 2004) to CCEWs (Kiladis et al. 2009) to intraseasonal oscillations on the planetary scale such as
38 the Madden-Julian Oscillation (MJO) (Zhang 2005). The early investigation about mean prop-
39 erties of tropical convection and its variability based on the GARP Atlantic Tropical Experiment
40 (GATE) dates back to 1970s (Houze Jr and Cheng 1977). Recently, organized tropical convec-
41 tion is documented in the Year of Tropical Convection (YOTC) virtual global field-campaign and
42 further analyzed through diagnostic, theoretical and numerical studies (Moncrieff et al. 2012). In
43 particular, superclusters of cloudiness and rainfall on the synoptic scale are frequently organized
44 by CCEWs that propagates eastward or westward along the equator or the intertropical Conver-
45 gence Zone (ITCZ) (Nakazawa 1988; Kiladis et al. 2009). For example, cloudiness variations
46 associated with convectively coupled Kelvin waves have a peak along the latitude of the climato-
47 logical ITCZ, indicating significant variability of convective activities (Wheeler et al. 2000; Ki-
48 ladis et al. 2009). The crucial impacts of convectively coupled Kelvin waves lie in their strong
49 interaction with the MJO (Straub et al. 2006) and linking synoptic-scale variation of the Atlantic
50 ITCZ with precipitation anomalies in South America (Wang and Fu 2007). Instead of organizing
51 in a single-scale convective envelope, CCEWs are manifested in a hierarchical structure where nu-
52 merous mesoscale convective elements are embedded in the synoptic-scale convective envelope.
53 The examples of CCEWs in such a multi-scale structure include the convectively coupled Kelvin
54 wave observed in the eastern Pacific ITCZ where many small-scale, westward-moving convective
55 elements over several hundred kilometers are found in the large-scale convective envelope (Straub
56 and Kiladis 2002), and the 2-day wave observed over the Indo-Pacific oceanic warm pool where

57 numerous embedded cloud clusters propagate at various speeds and directions (Chen et al. 1996).
58 These small-scale convective elements are categorized as MCSs, which are the dominant heavy
59 rain producers in the tropics and subtropics, providing more than 50% of the precipitation (Tao
60 and Moncrieff 2009).

61 The morphology of organized tropical convection is similar across multiple scales from the
62 mesoscale to synoptic scale to planetary scale, which is explained by a self-similarity princi-
63 ple derived by Majda (2007). One crucial feature of self-similarity is that both dynamical and
64 convective fields during tropical convection exhibit a front-to-rear vertical tilt. By regarding an
65 MCS as a small analogue or prototype of large-scale waves, the self-similarity of cloudiness is
66 explained as a similar progression from shallow to deep convection to stratiform anvils on many
67 time scales (Mapes et al. 2006). Such a trimodal characteristics of tropical convection including
68 cumulus congestus, deep convection and stratiform clouds is also found in a broad spectrum based
69 on shipboard radar data (Johnson et al. 1999). By parameterizing these three cloud types (con-
70 gestus, deep, stratiform) and carefully dealing with the transition between different type clouds,
71 the multcloud models successfully reproduce many key features of CCEWs including the spec-
72 trum peaks, reduced phase speed and self-similar front-to-rear tilt (Khouider and Majda 2008a).
73 Besides, the large-scale organization of tropical deep convection is investigated in idealized two-
74 dimensional cloud-resolving simulations (Grabowski and Moncrieff 2001). Their results highlight
75 westward-moving MCSs over several hundred kilometers embedded in an eastward-moving con-
76 vective envelope over several thousand kilometers.

77 Understanding the scale interaction between small-scale disturbances and large-scale wave en-
78 velope is crucial, not only for explaining propagation properties and spatial patterns of CCEWs,
79 but also improving the skill of global climate models (GCMs) for weather and climate forecast-
80 ing. Based on soundings, the momentum budget residual is estimated for the effects of convective

81 momentum transport (CMT) over the western Pacific warm pool (Tung and Yanai 2002a,b). In
82 general, CMT not only describes the momentum transport when organized moist convection on
83 smaller scales affects the large-scale flow field, but also involves the process of energy conversion
84 from convective available potential energy to horizontal kinetic energy. In the theoretical direc-
85 tions, simple stochastic models that capture the significant intermittent upscale impact of CMT
86 on the large scales due to organized unresolved convection from squall lines are built and further
87 tested in the column model environment and the organized synoptic-scale CCEWs through an
88 idealized multicloud model (Majda and Stechmann 2008). Besides, a simple dynamic model is
89 derived by including interactions between a large-scale zonal mean flow and convectively coupled
90 gravity waves and utilized to quantify and parameterize the effects of CMT (Majda and Stechmann
91 2009). Furthermore, CMT and its impact on the large-scale organization of convection are diag-
92 nostically investigated in the two-dimensional cloud-resolving model (Grabowski and Moncrieff
93 2001) and three-dimensional state-of-the-art mesoscale model (Khouider and Han 2013).

94 In spite of so much progress, the crucial features of the MCSs and their upscale impact on the
95 large-scale circulation and precipitation are still poorly simulated in the GCMs, which is mainly
96 related with the fact that the resolution of GCMs is too coarse to explicitly simulate the dynamical
97 and thermal properties of MCSs. In addition, there still exist huge discrepancies of precipitation
98 amounts between the comprehensive numerical simulations and the observed tropical convection.
99 For example, the present-day GCMs are still struggling to reproduce the realistic features of the
100 MJO (Jiang et al. 2015). One hypothesis for such huge discrepancies of precipitation amount is the
101 inadequate treatment of organized tropical convection and its missing upscale impact on the large-
102 scale flow field in the GCMs. The goals of this paper are as follows: first, using a simple multi-
103 scale model to capture the hierarchical structure of synoptic-scale equatorial waves with westward-
104 moving mesoscale systems embedded in a eastward-moving synoptic-scale convective envelope;

105 secondly, assessing the upscale impact of mesoscale disturbances on the behaviors of synoptic-
106 scale circulation through eddy transfer of momentum and temperature; thirdly, understanding how
107 much of synoptic-scale circulation is induced by eddy transfer of momentum and temperature
108 rather than synoptic-scale mean heating.

109 The simple multi-scale model used in this paper is the mesoscale equatorial synoptic dynamics
110 (MESD) model, originally derived by Majda (2007). In general, self-consistent multi-scale models
111 such as the MESD model were derived systematically by following multi-scale asymptotic theory
112 and used to describe the hierarchical structures of atmospheric flows in the tropics (Majda and
113 Klein 2003; Majda 2007; Yang and Majda 2014; Majda and Yang 2016). The advantages of using
114 these multi-scale models lie in isolating the essential components of multi-scale interaction and
115 providing assessment of the upscale impact of the small-scale fluctuations onto the large-scale
116 envelope through eddy flux divergence of momentum and temperature in a transparent fashion.
117 Specifically, the MESD model can be used to model the cluster-supercluster interactions across
118 mesoscale and synoptic scale and incorporate them together in a simple multi-scale framework.

119 In order to achieve the goals mentioned before, the two-dimensional MESD model is first set
120 up in the same manner as the two-dimensional cloud-resolving model (Grabowski and Mon-
121 crieff 2001). The resulting MESD model solutions are directly compared with those figures in
122 (Grabowski and Moncrieff 2001) in terms of zonal momentum, vertical velocity and potential tem-
123 perature anomalies. Then three different scenarios are discussed, where synoptic-scale circulation
124 is driven by various two-scale heating including elevated upright mean heating, faster propagating
125 convective envelope and westward-moving mesoscale heating in an upward/westward tilt. Sev-
126 eral crucial results are obtained by calculating eddy transfer of momentum and temperature and
127 comparing synoptic-scale circulation response to mean heating and eddy terms. First, the two-
128 dimensional MESD model successfully reproduces many key features of flow fields, including the

129 front-to-rear tilted vertical structure of mesoscale systems, and synoptic-scale circulation. Sec-
130 ondly, in the scenario with an upright mean heating, the total potential temperature anomalies and
131 zonal velocity still have front-to-rear tilted vertical structure, highlighting the significant upscale
132 impact of the mesoscale disturbances on the spatial pattern of synoptic-scale circulation. Thirdly,
133 in the scenario with a faster propagating convective envelope (less than 25 ms^{-1}), the MESD model
134 predicts that the synoptic-scale circulation response to the eddy terms become less important due
135 to the competing effects of eddy transfer of momentum and temperature, while that driven by
136 mean heating dominates. Such a result can explain the discrepancies between the cloud resolving
137 simulations in trade wind regime by Grabowski and Moncrieff (2001) where CMT is significant
138 and those in zero background mean flow regime by Tulich and Mapes (2008) where CMT is neg-
139 ligible. Lastly, in the scenario with westward-moving mesoscale heating in an upward/westward
140 tilt, positive potential temperature anomalies are induced by eddy terms in the leading edge of the
141 convective envelope, which tends to suppress shallow convection in a moist environment.

142 The rest of this paper is organized as follows. Sec.2 summarizes the properties of the two-
143 dimensional MESD model and self-similarity of flow field across mesoscale and synoptic scale.
144 Sec.3 discusses the prescribed westward-moving mesoscale heating, mesoscale fluctuations of
145 flow field and the associated eddy transfer of momentum and temperature. Sec.4 shows synoptic-
146 scale circulation response to the eastward-moving mean heating with embedded westward-moving
147 mesoscale heating, which are directly compared with the results from the cloud-resolving model
148 (Grabowski and Moncrieff 2001). Sec.5-7 consider three different scenarios with elevated upright
149 mean heating, faster propagating convective envelope and westward-moving mesoscale heating in
150 an upward/westward tilt. The paper ends with a concluding discussion. A proof about zero upscale
151 fluxes from free gravity waves is included in the Appendix.

152 **2. Properties of the MESD Model**

153 One common feature of CCEWs in the tropics is that numerous smaller-scale convective ele-
154 ments are embedded in the large-scale convective envelope (Straub and Kiladis 2002; Chen et al.
155 1996), which is a suitable scenario for using the multi-scale models. In general, the multispace,
156 multitime simplified asymptotic models are derived systematically from the equatorial primitive
157 equations on an equatorial β -plane, providing a useful framework to understand the multi-scale
158 phenomenon (Majda and Klein 2003; Majda 2007). Particularly, the three-dimensional MESD
159 model originally derived by Majda (2007) consists of two groups of linear primitive equations, one
160 of which governs irrotational mesoscale flows with gravity waves and the other one of which gov-
161 erns rotational synoptic-scale flows with baroclinic Kelvin waves, Rossby waves, mixed Rossby-
162 gravity waves, gravity waves and barotropic Rossby waves (Majda 2003). More importantly, eddy
163 transfer terms involving mesoscale fluctuations of momentum and temperature arise naturally in
164 the synoptic-scale equations and drive the synoptic-scale circulation response along with mean
165 heating, where the former is interpreted as the upscale impact of mesoscale disturbances on the
166 synoptic-scale circulation.

167 *a. The two-dimensional MESD model*

168 Rather than the original three-dimensional MESD model (Majda 2007), the two-dimensional
169 MESD model is used in this paper. Such a simplified version of MESD model can not only
170 simplify our discussion, but provide a suitable scenario to have direct comparison with the two-
171 dimensional cloud resolving model (Grabowski and Moncrieff 2001). In fact, it is quite straight-
172 forward to achieve such dimensionality reduction by just assuming that the dominant flow field
173 is near the equator (the Coriolis force is negligible) and meridionally symmetric (meridional ve-

174 locity and meridional momentum forcing vanishes). The resulting two-dimensional MESD model
 175 consists of two groups of equations on mesoscale and synoptic scale respectively.

176 The **equations for mesoscale dynamics** in dimensionless units read as follows,

$$u'_\tau = -p'_x + s'_u, \quad (1a)$$

$$\theta'_\tau + w' = s'_\theta, \quad (1b)$$

$$p'_z = \theta', \quad (1c)$$

$$u'_x + w'_z = 0, \quad (1d)$$

177 where the prime means mesoscale fluctuations of flow fields. s'_u and s'_θ stand for momentum and
 178 thermal forcing on the mesoscale.

179 The **equations for synoptic-scale dynamics** in dimensionless units read as follows,

$$U_t = -P_X - dU - \langle \overline{w'u'} \rangle_z + \langle \bar{S}^u \rangle, \quad (2a)$$

$$\Theta_t + W = -d_\theta \Theta - \langle \overline{w'\theta'} \rangle_z + \langle \bar{S}^\theta \rangle, \quad (2b)$$

$$P_z = \Theta, \quad (2c)$$

$$U_X + W_z = 0, \quad (2d)$$

180 where $\langle \bar{S}^u \rangle$ and $\langle \bar{S}^\theta \rangle$ stand for momentum and thermal forcing on the synoptic scale.

181 A linear momentum damping term is added in Eqs.2a to mimic momentum dissipation of cu-
 182 mulus drag. The coefficient d in units of $1/day$ sets the time scale for momentum dissipation.

183 According to the observation, momentum damping time scale at the surface of the Pacific ocean
 184 could be as strong as 1 day (Deser 1993) while that at the upper troposphere is much longer. In

185 general, the momentum damping of large-scale circulation occurs on a time scale of $\mathcal{O}(1 - 10)$

186 days, and also depends on the vertical wavelength of the wind profile (Romps 2014). Besides, the

187 Newtonian cooling term $-d_\theta \Theta$ is added in the thermal equation in Eq.2b to mimic radiative cool-

188 ing. Such thermal damping in two-dimensional MESD model is necessary, otherwise potential
 189 temperature anomalies induced by heating forcing in Eq.2b grows linearly to infinity.

190 The two-dimensional MESD model describes hierarchical structure of tropical flows across mul-
 191 tiple spatial and temporal scales. One dimensionless unit of x, τ corresponds to 150 *km*, 50 *min*
 192 on the mesoscale, while that of X, t corresponds to 1500 *km* and 8.3 *h* on the synoptic scale. On
 193 these two scales, one dimensionless unit of some physical variables corresponds to the same di-
 194 mensional value, including zonal velocity u, U (5 ms^{-1}), pressure perturbation p, P ($250 \text{ m}^2 \text{ s}^{-2}$),
 195 potential temperature anomalies θ, Θ (3.3 K). In contrast, vertical velocity in one dimensionless
 196 value corresponds to 0.16 ms^{-1} on the mesoscale and 0.016 ms^{-1} on the synoptic scale. The
 197 vertical coordinate z is shared by groups of equations on different scales. Furthermore, the mo-
 198 mentum and thermal forcing on the synoptic scale is much weaker than those on the mesoscale by
 199 one order, which is consistent with the observation that the measured atmospheric heating from
 200 latent heat release in the tropics is in general much weaker on synoptic and even larger scales
 201 (Biello and Majda 2010). More details for all constants and physical parameters are summarized
 202 in Table.1.

203 Eqs.2a-2d involve mesoscale zonal and temporal averaging operators defined as follows (f is an
 204 arbitrary function),

$$\bar{f}(X, z, t, \tau) = \lim_{L \rightarrow \infty} \frac{1}{2L_x} \int_{-L_x}^{L_x} f(X, x, z, t, \tau) dx \quad (3)$$

$$\langle f \rangle(X, x, z, t) = \lim_{T \rightarrow \infty} \frac{1}{2T} \int_{-T}^T f(X, x, z, t, \tau) d\tau \quad (4)$$

205 where L_x is the mesoscale zonal length of the domain and T is the time range in the asymptotic
 206 limit. In Eqs.1a-1d, all mesoscale physical variables f' satisfy $\bar{f}' = 0$ and $\langle f' \rangle = 0$, representing
 207 mesoscale fluctuations of flow fields.

208 *b. Self-similarity of flow fields across mesoscale and synoptic scale*

209 The governing equations for mesoscale dynamics in Eqs.1a-1d and synoptic-scale dynamics in
 210 Eqs.2a-2d have many features in common. Specifically, if the momentum damping term $-dU$,
 211 radiative cooling term $-d_\theta\Theta$ and all eddy terms $-\langle w'u' \rangle_z$, $-\langle w'\theta' \rangle_z$ are merged into the general
 212 forcing terms on the right hand side, the mesoscale dynamics in Eqs.1a-1d are exactly the same
 213 as the synoptic-scale dynamics in Eqs.2a-2d. Such self-similarity principle is derived by Majda
 214 (2007).

215 With rigid-lid boundary conditions imposed, the governing equations in common across
 216 mesoscale and synoptic scale have explicit solution formulas involving the barotropic mode and an
 217 infinite set of baroclinic modes (Majda 2003). For vertical decomposition, all physical variables
 218 are expanded into different vertical modes with sine and cosine functions as follows,

$$f = \sum_{q=0}^{\infty} f_q \cos(qz), f \in \{u, p, s^u\} \quad (5)$$

$$g = \sum_{q=1}^{\infty} g_q [-q \sin(qz)], g \in \{\theta, w, s^\theta\} \quad (6)$$

219 where q is the index for different vertical modes. The barotropic mode in Eq.6 vanishes due to the
 220 rigid-lid boundary condition.

221 After plugging all physical variables into the ansatz in Eqs.5-6, the governing equations are
 222 decomposed into different groups of equations in barotropic and baroclinic modes. The barotropic
 223 mode ($q = 0$) is governed by the following equations,

$$u_t = -p_x + s^u \quad (7a)$$

$$u_x = 0 \quad (7b)$$

224 which states that the time tendency of mean zonal velocity \bar{u}_t is only forced by mean zonal momen-
 225 tum forcing \bar{s}'' , while nonzero wavenumber mode of zonal velocity vanishes and that of pressure
 226 gradient p_x is balanced by zonal momentum forcing s'' .

227 The baroclinic mode ($q \neq 0$) is governed by the following equations,

$$u_t = -p_x + s'' \quad (8a)$$

$$\theta_t + w = s^\theta \quad (8b)$$

$$p = \theta \quad (8c)$$

$$u_x - q^2 w = 0 \quad (8d)$$

228 where q is the vertical index. The subscripts and superscripts of physical variables are ignored
 229 for simplicity. After plugging in the plane wave ansatz, the dispersion relation of gravity waves is
 230 obtained,

$$\omega = \pm \frac{k}{q} \quad (9)$$

231 which indicates that the phase speed in the q baroclinic mode is $c = \frac{\omega}{k} = \pm \frac{1}{q}$ in both eastward
 232 and westward directions. Also, gravity waves in higher baroclinic modes propagate slower. For
 233 example, in dimensional units, the phase speeds of gravity waves in the 1st, 2nd and 3rd baroclinic
 234 modes are 50 ms^{-1} , 25 ms^{-1} and 16.7 ms^{-1} .

235 3. Westward-Moving Mesoscale Disturbances and the Upscale Flux

236 Inside the convective envelope of synoptic-scale waves, numerous convective elements on the
 237 smaller scales are found, which is associated with MCSs such as squall-line systems (Houze Jr
 238 1975, 1977; Houze Jr and Cheng 1977; Houze 2004). Such hierarchical structure of the synoptic-
 239 scale convective envelope with embedded MCSs is captured in an idealized two-dimensional cloud
 240 resolving simulations (Grabowski and Moncrieff 2001). Within the large-scale organization of

241 convection, numerous westward-moving systems are identified, which are characterized by low-
 242 level inflow and upper-level outflow to both east and west with extensive stratiform cloud in the
 243 upper troposphere.

244 The main goal of this section is to model such mesoscale convective systems by using the
 245 mesoscale dynamics equations in Eqs.1a-1d and prescribing top-heavy diabatic heating to mimic
 246 latent heat release. In the vertical direction, the rigid-lid boundary conditions are imposed,

$$w = 0, \text{ at } z = 0, \pi. \quad (10)$$

247 where $z = 0, \pi$ denote the surface and the top of troposphere. The mesoscale solutions are assumed
 248 to be periodic in the zonal direction. The vertical extent of the domain is 15.7 km . Since it is proved
 249 in the Appendix that free gravity waves on the mesoscale can not generate upscale fluxes, here the
 250 mesoscale solutions forced by prescribed heating are solved analytically and all free gravity waves
 251 are ignored.

252 *a. Westward-moving mesoscale heating*

253 In order to mimic latent heat release from westward-moving systems with extensive stratiform
 254 clouds in the upper troposphere in (Grabowski and Moncrieff 2001), a top-heavy mesoscale heat-
 255 ing is prescribed below,

$$s'_{\theta} = c_0 [\sin(kx + \omega\tau) \sin(z) + \alpha \sin(kx + \omega\tau + \phi_0) \sin(2z)] \quad (11)$$

256 where the magnitude coefficient is $c_0 = 0.7$ in dimensional unit 100 Kday^{-1} . The wavenumber
 257 $k = \frac{3\pi}{5}$ is chosen for wavelength 500 km , qualitatively consistent with typical length scales of MCSs
 258 (Houze 2004). The frequency $\omega = \frac{12}{50}k$ is picked so that the associated phase speed reaches 12
 259 ms^{-1} , the same as the averaged westward propagation speed of mesoscale systems in (Grabowski
 260 and Moncrieff 2001). The strength coefficient of the second baroclinic mode α is set as -0.9 to

261 obtain top-heavy heating profile. The phase lag $\phi_0 = -\pi/4$ between the first and second baroclinic
262 modes is used to mimic the upward/eastward tilted heating.

263 Fig.1 shows the prescribed mesoscale heating profile in the longitude-height diagram. Such
264 a mesoscale heating in a front-to-rear tilt is frequently observed in MCSs (Houze Jr 1977;
265 Houze 2004). Meanwhile, heating and cooling regions reach their maximum magnitudes about
266 $100Kday^{-1}$ in the upper troposphere at the height 11.5 km. In fact, heating profiles with the first
267 baroclinic mode for deep convective heating and the second baroclinic mode for both congestus
268 and stratiform heating have been utilized successfully in the multcloud models to reproduce many
269 realistic features of tropical convection (Khouider and Majda 2006b,a, 2007).

270 *b. Mesoscale fluctuations of zonal velocity and potential temperature anomalies*

271 Fig.2a shows the solution of zonal velocity in the longitude-height diagram, which is directly
272 compared with Fig.2b from the study (Grabowski and Moncrieff 2001). These two panels are
273 qualitatively similar to each other and both feature zonal velocity in a front-to-rear tilt. In details,
274 zonal velocity in this westward-moving mesoscale system is characterized by an inflow layer of
275 westerly winds, which starts from the lower levels to the west, lifts up to the middle troposphere
276 in the middle of the domain and continuously extends to the upper troposphere to the east. The
277 maximum magnitude of westerly winds is reached in the upper troposphere to the east. Besides,
278 there are low-level easterly winds to the east and upper-level easterly winds to the west. Fig.2c
279 shows the stream lines of mesoscale flow fields. When compared with Fig.2d from the study
280 (Grabowski and Moncrieff 2001), both panels are characterized by wind convergence in the lower
281 troposphere and wind divergence in the upper troposphere in a front-to-rear tilt.

282 Fig.2e shows potential temperature anomalies in the longitude-height diagram. In the vertical
283 direction, potential temperature anomalies are mostly characterized by opposite anomalies be-

284 tween upper and lower tropospheres, indicating the significant magnitude of the second baroclinic
 285 mode. Fig.2f shows vertical velocity in the longitude-height diagram, whose spatial pattern is
 286 quite similar to that of mesoscale heating. In connection with mesoscale heating in Fig.1, heating
 287 regions are dominated by the westward/downward winds while cooling regions are dominated by
 288 the eastward/upward winds.

289 *c. Eddy transfer of momentum and temperature*

290 In the MESD model, the governing equations for synoptic-scale dynamics is forced by two eddy
 291 terms on the right hand side in Eq.2a-2b. By utilizing the mesoscale solutions of zonal and vertical
 292 velocity, eddy momentum transfer (EMT) has the following expression,

$$\begin{aligned}
 F^u &= - \langle \overline{w'u'} \rangle_z \\
 &= \kappa^u \left[-\frac{3}{2} \cos(z) + \frac{3}{2} \cos(3z) \right]
 \end{aligned}
 \tag{12}$$

293 which represents upscale impact of mesoscale disturbances on the synoptic-scale circulation in the
 294 zonal momentum budget. In fact, EMT is referred as CMT, which has been studied from different
 295 perspectives to highlight its significance such as stochastic models (Majda and Stechmann 2008;
 296 Khouider et al. 2012) and dynamical models with cloud parameterization (Majda and Stechmann
 297 2009). Physically, positive (negative) anomalies of EMT corresponds to eastward (westward)
 298 momentum forcing on the synoptic-scale circulation. One crucial feature of the EMT term is that
 299 its vertical profile is independent of all the parameters, while its magnitude and directions are
 300 determined by the coefficient κ^u in the following explicit expression,

$$\kappa^u = \frac{\sin(\phi_0) \alpha k^3}{2(\omega^2 - k^2)(4\omega^2 - k^2)}.
 \tag{13}$$

301 Since the product term $\sin(\phi_0)\alpha$ controls the vertical structure of mesoscale heating in Eq.11,
 302 its tilting direction (upward/eastward or upward/westward) determines the sign of EMT. Besides,

303 there exist two critical phase speeds of mesoscale heating $|\frac{\omega}{k}| = 1, 0.5$, which actually corresponds
 304 to those of the first and second baroclinic modes as discussed in Sec.2b.

305 Fig.3a shows the vertical profile of eddy fluxes of zonal momentum and EMT. The eddy fluxes of
 306 zonal momentum reaches its maximum positive value in the middle troposphere and decays to zero
 307 in both the upper and lower boundaries, due to the in-phase relation between zonal velocity and
 308 vertical velocity as shown in Fig.2. The EMT term is characterized by its positive anomalies in the
 309 upper troposphere near the height 11 *km* and negative anomalies in the lower troposphere near the
 310 height 5 *km*. Such upper-level eastward momentum forcing and low-level westward momentum
 311 forcing represent vertical transfer of zonal momentum on the synoptic scale, resulting in vertical
 312 shear of zonal winds.

313 The other eddy term that appears at the right hand side of the synoptic-scale equations is eddy
 314 heat transfer (EHT). By utilizing the mesoscale solutions of vertical velocity and potential temper-
 315 ature anomalies, its expression is as follows,

$$\begin{aligned}
 F^\theta &= -\langle w'\theta' \rangle_z \\
 &= \kappa^\theta \left[-\frac{3}{2} \sin(z) + \frac{9}{2} \sin(3z) \right]
 \end{aligned} \tag{14}$$

316 which represents upscale impact of mesoscale disturbances on the synoptic-scale thermal dynam-
 317 ics. Physically, positive (negative) anomalies of EHT corresponds to heating (cooling) on the
 318 synoptic-scale circulation. The coefficient κ^θ ,

$$\kappa^\theta = \frac{\sin(\phi_0) \alpha k^2 \omega}{2(\omega^2 - k^2)(4\omega^2 - k^2)}. \tag{15}$$

319 which is quite similar to that in Eq.13 except that wavenumber k in the numerator is replaced by
 320 frequency ω . Suppose wavenumber k is fixed, the sign of frequency ω determines the propagation
 321 direction of mesoscale heating. The product term $\sin(\phi_0)\alpha$ determines the front-to-rear tilted

322 vertical structure. Therefore, the sign of EHT is determined by the tilting direction of mesoscale
323 heating, compared with its propagation direction.

324 Fig.3b shows the vertical structure of eddy fluxes of temperature and EHT. The eddy fluxes of
325 temperature reaches positive value in the upper troposphere near the height 11 *km* and negative
326 value in the lower troposphere near the height 5 *km*, due to the opposite phase relation between
327 potential temperature anomalies and vertical velocity in upper and lower tropospheres as shown in
328 Fig.2. The EHT term is characterized by heating in both the lower and upper troposphere and cool-
329 ing in the middle troposphere. In a moist environment, the heating in the lower troposphere could
330 suppress the convection through increasing saturation rate of vapor and convective inhibition.

331 The ratio between EMT and EHT in dimensionless units determines relative strength of the
332 corresponding synoptic-scale circulation response. In fact, the ratio between the magnitude coef-
333 ficients κ^u and κ^θ is equal to,

$$\frac{\kappa^\theta}{\kappa^u} = \frac{\omega}{k}, \quad (16)$$

334 which is equal to the phase speed of mesoscale heating and suggests that the strength of the corre-
335 sponding circulation response to EMT and EHT depends on the propagation speed of the diabatic
336 heating. Suppose the propagation speed of the diabatic heating is relatively slow, the synoptic-
337 scale circulation response due to the upscale impact of mesoscale disturbances is mostly induced
338 by the EMT. While in the fast propagating mesoscale heating scenario, the synoptic-scale circula-
339 tion response is mostly induced by the EHT.

340 **4. Synoptic-Scale Circulation Response to Upscale Impact of Mesoscale Disturbances and**

341 **Mean Heating**

342 The hierarchical structure of tropical convection has been identified by Nakazawa (1988) and
343 further explained as a synoptic-scale eastward-moving convective envelope in a supercluster em-

344 bedded by mesoscale westward-moving disturbances in cloud clusters. As discussed in Sec.3,
 345 mesoscale fluctuations in tilted vertical structure tend to generate eddy transfer of momentum and
 346 temperature and drive synoptic-scale circulation response. In this section, the synoptic-scale circu-
 347 lation response induced by upscale impact of mesoscale disturbances and mean heating is directly
 348 compared with results from the study (Grabowski and Moncrieff 2001).

349 Here the equations for synoptic-scale dynamics in Eqs.2a-2d are used. By ignoring the synoptic-
 350 scale momentum forcing in Eq.2a, the synoptic-scale circulation response is driven by mean heat-
 351 ing , EMT and EHT. As for boundary conditions, the solutions are assumed to be periodic in the
 352 zonal direction. In the vertical direction, rigid-lid boundary conditions are imposed,

$$W = 0, \text{ at } z = 0, \pi \quad (17)$$

353 where $z = 0, \pi$ stand for surface and top of the troposphere. The full domain is $0 \leq x < 20,000km$,
 354 $0 \leq z \leq 15.7km$. As for the damping terms, the momentum damping coefficient $d = d(z)$ is as-
 355 sumed to be a linear function of height, which contains 16.6 *h* damping time scale at surface and
 356 2.9 *day* damping time scale at top of the troposphere. The radiative cooling coefficient $d_\theta = 0.1572$
 357 (2.2 *day*) is assumed to be homogeneous throughout the whole domain, similar to (Grabowski and
 358 Moncrieff 2001). All physical variables are initialized from a state of rest and run for 41.5 *day*.
 359 The details for spatial and temporal resolutions are summarized in Table.2.

360 Tropical cloud clusters within the superclusters are found to experience different life stages
 361 from the developing to the east of the convective envelope to the mature stage in the middle to
 362 the decaying stage to the west (Nakazawa 1988). The large-scale modulation of the mesoscale
 363 fluctuations is represented by a synoptic-scale envelope function below,

$$E(X - st) = 0.5 + 0.5 \sin\left(\frac{10\pi}{L}(X - st)\right) \quad (18)$$

364 where $s = \frac{4}{25}$ (corresponds to 8 ms^{-1}) is the propagation speed of the envelope. $L = 40/3$
 365 (20,000km) is the zonal extent of the whole domain. The envelope function has 5 complete peri-
 366 ods throughout the whole domain, the same as (Grabowski and Moncrieff 2001). The mesoscale
 367 heating in a synoptic-scale convective envelope is reformulated as follows,

$$s'_\theta = E(X - st) c_0 [\sin(kx + \omega\tau) \sin(z) + \alpha \sin(kx + \omega\tau + \phi_0) \sin(2z)], \quad (19)$$

368 where all the parameters for the mesoscale heating are the same as Eq.11.

369 *a. Domain-averaged zonal velocity*

370 According to Eq.2a, the domain-averaged zonal velocity is governed by,

$$\frac{\partial \bar{U}}{\partial t} = -d\bar{U} + \bar{F}^u, \quad (20)$$

371 where the long bar denotes zonal averaging about synoptic-scale X . Eq.20 indicates that the
 372 domain-averaged zonal velocity is only related with momentum damping coefficient d and EMT
 373 F^u , while independent of EHT F^θ and mean heating $\langle \bar{S}^\theta \rangle$. In order to discuss domain-averaged
 374 zonal velocity, we only need to consider synoptic-scale circulation response to EMT.

375 The momentum flux $w'u'$ describes vertical transport of zonal momentum on the synoptic scale.
 376 As shown by Fig.4b, the momentum flux in the study (Grabowski and Moncrieff 2001) is char-
 377 acterized by five nearly equally spaced anomalies with their maximum value in the middle tro-
 378 posphere. Such mid-level momentum fluxes are captured in the numerical solutions in Fig.4a.
 379 According to Fig.3a, momentum fluxes reach their maximum value in the middle troposphere,
 380 which is consistent with the in-phase zonal and vertical velocity fields in Sec.3. Here the maxi-
 381 mum magnitude of momentum fluxes is $0.18 \text{ m}^2\text{s}^{-2}$, which qualitatively matches that in the study
 382 (Grabowski and Moncrieff 2001). As modulated by the envelope function in Eq.18, the momen-
 383 tum flux also has five positive anomalies along the whole domain.

384 Fig.4 shows zonal velocity induced by EMT in the longitude-height diagram at 41.5 *day*. The
385 vertical profile of zonal velocity is characterized by low-level easterly winds at the height $z = 5km$
386 and upper-level westerly winds at the height $z = 11km$. Such vertical shear of zonal winds is
387 consistent with that of EMT in Fig.3a. Due to the vertically decaying momentum damping coeffi-
388 cient, the maximum strength of zonal velocity in the upper troposphere is double as much as that
389 in the lower troposphere. Near the surface and top of the troposphere, there are alternate easterlies
390 and westerlies in weak magnitude. As shown by Fig.4e, the domain-averaged zonal velocity in
391 the study (Grabowski and Moncrieff 2001) is characterized by strong westerly winds ($2ms^{-1}$) in
392 the upper troposphere and relatively weak easterly winds ($1ms^{-1}$) in the lower troposphere. Such
393 vertical shear of zonal velocity and comparable magnitude is captured in the MESD model nu-
394 merical solutions as shown in Fig.4d. According to Eq.20, in the steady state, EMT is balanced
395 by the zonal momentum damping, resulting in a vertical shear of zonal velocity. Such wind shear
396 has been recognized to play a critical role for tropical convection and its organization (Grabowski
397 et al. 1996; Grabowski and Moncrieff 2001).

398 *b. Potential temperature anomalies on the synoptic scale*

399 The spatial distribution of temperature perturbations in the study (Grabowski and Moncrieff
400 2001) is shown in Fig.6f, which is characterized by anomalies in a front-to-rear tilt. In detail, the
401 temperature anomalies have an elbow shape with upward/westward tilted structure below the level
402 $z = 14km$ and upward/eastward tilted structure above that level, which is reminiscent of the typical
403 temperature anomalies associated with CCEWs as observed in nature (Kiladis et al. 2009). The
404 maximum temperature perturbations are about $0.8 K$ and the minimum temperature perturbations
405 are about $-0.6 K$, both of which are located in the upper troposphere. Such potential temperature
406 anomalies can be either driven by mean heating or eddy transfer of momentum and temperature,

407 or both of them. To figure out this question, potential temperature anomalies induced by mean
 408 heating, EMT and EHT are discussed separately and the total anomalies are compared directly
 409 with those from the study (Grabowski and Moncrieff 2001).

410 Physically, latent heat release during tropical precipitation starts from the low-level congestus
 411 clouds to deep convective clouds to upper-level stratiform clouds (Khouider and Majda 2008b).
 412 Thus mean heating is prescribed as follows,

$$\langle \bar{S}^\theta \rangle = c_0 \left[\sin \left(\frac{10\pi(X-st)}{L} \right) \sin(z) - \frac{2}{3} \sin \left(\frac{10\pi(X-st)}{L} + \frac{\pi}{2} \right) \sin(2z) \right], \quad (21)$$

413 where magnitude coefficient $c_0 = 0.5$ corresponds to 5 Kday^{-1} . $s = \frac{4}{25}$ (8 ms^{-1}) is the propagation
 414 speed of the envelope. Eq.21 also assumes that the convective envelope $E(X-st)$ in Eq.18 is in
 415 phase with the first baroclinic mode (deep convection) of mean heating.

416 Fig.5 shows the prescribed mean heating and the resulting zonal/vertical velocity in the
 417 longitude-height diagram at 41.5 *day*. The alternate heating and cooling regions are up-
 418 ward/westward tilted. Upward/westward motion dominates in heating regions, while down-
 419 ward/eastward motion dominates in cooling regions. Such upward/westward tilted vertical struc-
 420 ture of zonal velocity is quite often observed in the reality (Kiladis et al. 2009). As for the vertical
 421 velocity, the upward (downward) motion is in phase with the heating (cooling) regions, which is
 422 reminiscent of the realistic process in nature when air parcels gain buoyancy and lift up during
 423 latent heat release. Besides, wind convergence (divergence) occurs in the leading edge of heating
 424 (cooling) regions and trailing edge of cooling (heating) regions. In a moist environment, such
 425 wind convergence in the leading edge tends to trap moisture and bring them along with the inflow
 426 to the upper levels.

427 Fig.6a shows potential temperature anomalies induced by mean heating in the longitude-height
 428 diagram at 41.5 *day*. Again, the spatial pattern of potential temperature anomalies is characterized

429 by the upward/westward tilted vertical structure. Besides, opposite potential temperature anoma-
430 lies are induced in the upper and lower tropospheres, indicating the significant synoptic-scale
431 circulation response in the second baroclinic mode. Also, the maximum magnitude of potential
432 temperature anomalies is reached in the lower troposphere at the height 4 *km* and upper tropo-
433 sphere at the height 12 *km*. Such potential temperature anomalies are quite different from those in
434 the study (Grabowski and Moncrieff 2001).

435 Fig.6b potential temperature anomalies induced by EMT in the longitude-height diagram at
436 41.5 *day*. The vertical profile of potential temperature anomalies is characterized by positive
437 (negative) anomalies in the middle troposphere and negative (positive) anomalies in the upper and
438 lower tropospheres, indicating the significant strength of the third baroclinic mode. Fig.6d shows
439 potential temperature anomalies induced by EHT in the longitude-height diagram at 41.5 *day*.
440 The maximum magnitude of potential temperature anomalies induced by EHT is much weaker
441 than that by EMT, which can be explained by the slow phase speed of mesoscale heating in Eq.16.
442 Besides, potential temperature anomalies induced by EMT and EHT are almost opposite to each
443 other, resulting in weaker magnitude of total potential temperature anomalies.

444 Fig.6c shows relative location between potential temperature anomalies induced by eddy terms
445 and those induced by mean heating in the longitude-height diagram at 41.5 *day*. In the lower
446 troposphere, the anomalies induced by eddy terms and mean heating overlap each other but in
447 the opposite signs, resulting in diminishing potential temperature anomalies below the height 5
448 *km*. In fact, such diminishing potential temperature anomalies is apparent in Fig.6f from the study
449 (Grabowski and Moncrieff 2001). In the middle troposphere, potential temperature anomalies
450 induced by eddy terms tends to strengthen anomalies induced by mean heating. In the upper
451 troposphere, potential temperature anomalies induced by eddy terms are out of phase with those
452 anomalies induced by mean heating, introducing the vertical shear of total potential temperature

453 anomalies and contributing the upward/westward tilted structure. As shown by Fig.6e, the total
454 potential temperatures anomalies are quite similar to those in Fig.6f from the study (Grabowski
455 and Moncrieff 2001).

456 *c. Zonal velocity on the synoptic scale*

457 Fig.7f shows the spatial distribution of zonal velocity in the study (Grabowski and Moncrieff
458 2001). In the lower (upper) troposphere, there are weak easterly (westerly) wind anomalies in
459 an upward/westward tilt. Besides, easterly wind anomalies exist between the strong westerly
460 wind anomalies near the top of troposphere. Without changing any model setup and physical
461 parameters, the goal of this section is to explore whether the total zonal velocity induced by mean
462 heating and eddy terms owns some similar features as Fig.7f.

463 Fig.7a shows zonal velocity induced by mean heating in the longitude-height diagram at 41.5
464 *day*. In connection with Fig.5, heating regions are mostly dominated by easterly wind anomalies,
465 while cooling regions are mostly dominated by westerly wind anomalies. At the surface, there
466 are easterly (westerly) wind anomalies to the east (west) of the heating regions, resulting in wind
467 convergence. Such upward/westward tilted zonal velocity features the significant magnitude of the
468 first and second baroclinic modes. When compared with Fig.7f, there exist many discrepancies of
469 zonal velocity anomalies such as the strong westerlies in the lower troposphere.

470 Fig.7b shows the zonal velocity induced by EMT in the longitude-height diagram at 41.5 *day*,
471 which is characterized by low-level easterly winds and upper-level westerly winds. Near the ver-
472 tical boundaries, westerly (easterly) wind anomalies in weak magnitude are also induced at the
473 surface (top). Fig.7d shows zonal velocity induced by EHT in the longitude-height diagram at
474 41.5 *day*. The maximum magnitude of zonal velocity is much weaker than that in Fig.7b. The
475 total zonal velocity induced by EMT and EHT is shown in Fig.7c, whose the maximum upper-

476 level westerlies and low-level easterlies are displaced to the west, due to zonal velocity anomalies
477 induced by EHT in the opposite sign.

478 Fig.7c shows relative location between the zonal velocity induced by mean heating and that
479 induced by eddy terms in the longitude-height diagram at 41.5 *day*. In the upper troposphere
480 between the height 8 *km* and 13 *km*, westerly wind anomalies induced by eddy terms dominates,
481 which tends to strengthen westerlies and weaken easterlies induced by mean heating. The resulting
482 total westerly wind anomalies in Fig.7e have upward/westward tilted vertical structure in the upper
483 troposphere. Similarly, easterly wind anomalies induced by eddy terms dominate in the lower
484 troposphere and the resulting total easterly wind anomalies have upward/westward tilted vertical
485 structure. Near the upper boundary, there also exist alternate easterly and westerly wind anomalies.
486 All these features in the upper and lower tropospheres resemble those in Fig.7f, except for the extra
487 westerly wind anomalies at the surface. One possible reason for such discrepancy is the missing
488 boundary layer dynamics in the MESD model but well resolved in the cloud resolving model
489 (Grabowski and Moncrieff 2001).

490 *d. Vertical velocity on the synoptic scale*

491 Fig.8d shows the spatial distribution of vertical velocity in the study (Grabowski and Moncrieff
492 2001). The most significant vertical motion is located in the middle and upper tropospheres with
493 the maximum value at height 12 *km*. In contrast to potential temperature anomalies and zonal ve-
494 locity in an upward/westward tilt, the vertical profile of vertical motion is mostly upright. Without
495 changing any model setup and physical parameters, the goal of this section is to explore whether
496 the total vertical velocity induced by mean heating and eddy terms has some similar features as
497 Fig.8f.

498 Fig.8a-b show vertical velocity induced by EMT and EHT in the longitude-height diagram at
499 41.5 *day*. Particularly, vertical velocity induced by eddy terms has very weak magnitude and
500 cancels each other in the opposite sign. In contrast, vertical velocity induced by mean heating
501 dominates in a comparable magnitude as Fig.8d. However, such front-to-rear tilted vertical motion
502 is quite different from the upright vertical motion in the study (Grabowski and Moncrieff 2001).
503 The next section is used to investigate whether upright mean heating will induce such upright
504 vertical motion and whether the upward/westward tilted vertical structure of potential temperature
505 anomalies and zonal velocity in the previous discussion is still captured in the existence of upright
506 mean heating and eddy terms.

507 **5. Upright Mean Heating**

508 According to the observation (Straub and Kiladis 2002; Kiladis et al. 2009), the dynamical fields
509 such as the zonal wind, temperature and specific humidity of convectively coupled Kelvin waves
510 have the upward/westward tilted vertical structure as they propagate eastward. Three cloud types
511 with low-level congestus clouds in the leading edge, deep convective clouds in the middle and
512 upper-level stratiform clouds in the trailing edge provides the key components for the tilted verti-
513 cal structure. However, Fig.8 suggests that the upright mean heating can induce upright vertical
514 motion on the large-scale domain. In this section, an elevated upright mean heating is used to
515 investigate whether both upright vertical velocity and upward/westward tilted zonal velocity and
516 potential temperature anomalies can still be captured in the existence of upright mean heating and
517 eddy terms.

518 In the tropics, deep convective clouds can warm and dry the entire troposphere through large
519 amounts of rainfall. Besides, the stratiform clouds warm and dry the upper troposphere through
520 stratiform precipitation and cool and moisten the lower troposphere due to the rain evaporation

521 (Khouider and Majda 2008b). Thus the latent heat release associated with in-phase deep and
522 stratiform convection is characterized by an elevated upright mean heating. In fact, it has been
523 confirmed by model experiments that diabatic processes in deep convective and stratiform regions
524 are essential to the formation of multiscale convective wave patterns (Tulich and Mapes 2008).
525 Here the vertical profile of the elevated upright mean heating is prescribed as follows,

$$G(z) = c_0 [\sin(z) - 0.6 \sin(2z)] \quad (22)$$

526 where the heating magnitude $c_0 = 0.5$ corresponds to 5 Kday^{-1} .

527 Fig.9 shows the elevated upright mean heating in the longitude-height diagram at 41.5 day .
528 The maximum magnitude of heating and cooling is achieved in the upper troposphere around 6
529 Kday^{-1} . Fig.9b shows vertical velocity induced by such an elevated upright mean heating in the
530 longitude-height diagram at 41.5 day . Similar to the mean heating, the vertical motion reaches
531 its maximum value in the upper troposphere at the height 12 km and decays as the height goes
532 close to the top and lower troposphere. More importantly, it is quite similar to Fig.8d in the study
533 (Grabowski and Moncrieff 2001), providing a convincing evidence for the upright mean heating
534 on the synoptic scale.

535 Fig.9c shows potential temperature anomalies induced by upright mean heating and eddy terms
536 in the longitude-height diagram at 41.5 day . Similar to those induced by upward/westward tilted
537 mean heating in Fig.6a, potential temperature anomalies induced by upright mean heating are also
538 characterized by opposite anomalies between the upper and lower tropospheres, indicating the
539 significant magnitude of the second baroclinic mode. Instead of an upward/west tilt, their vertical
540 profile has an upward/eastward tilt. After combined with those induced by eddy terms, the total
541 potential temperature anomalies as shown in Fig.9d exhibit an upward/westward tilt again with

542 the maximum value in the height 10 km. Such potential temperature anomalies are quite similar to
543 those in Fig.6f, except that the magnitude of upper-level anomalies here is much weaker.

544 Fig.9e shows zonal velocity induced by upright mean heating and eddy terms in the longitude-
545 height diagram at 41.5 day. In the heating region near the longitude $10^4 km$, there are wind con-
546 vergence in the middle troposphere at the height 7.85 km and wind divergence at the top of tro-
547 posphere. The wind strength at lower levels is negligible, due to the elevated height of mean
548 heating. After combined with that induced by eddy terms, the total zonal velocity as shown in
549 Fig.9f is characterized by upper-level westerlies on top of low-level easterlies. At the top of the
550 troposphere, there exist alternate easterlies and westerlies in strong magnitudes. The overall spa-
551 tial pattern of zonal velocity is quite similar to that in Fig.7f, except for the upper-level westerlies
552 in both upward/eastward and upward/westward tilts.

553 **6. Eastward-Moving Convective Envelope in Faster Propagation Speeds**

554 In the classic shallow water theory, the eastward-moving Kelvin waves in the first baroclinic
555 mode has phase speed $50 ms^{-1}$ (Matsuno 1966; Majda 2003). Because of moist processes,
556 CCEWs such as convectively coupled Kelvin waves typically have shallower equivalent depth
557 (slower phase speed). In general, the eastward-propagating convectively coupled Kelvin waves
558 are observed to have phase speed $15 - 20ms^{-1}$ over the west Pacific and $12 - 15ms^{-1}$ over the
559 Indian Ocean (Kiladis et al. 2009). Such slow propagation speeds of CCEWs become a bench-
560 mark to examine the behavior and skill of complex numerical simulations. In the two-dimensional
561 cloud resolving simulations with a strong background easterlies winds ($-10 ms^{-1}$) (Grabowski
562 and Moncrieff 2001), the large-scale envelope over a few thousand kilometers propagates west
563 to east at about 6 to $8 ms^{-1}$ relative to the earth, which is much slower than the phase speed of
564 the convectively coupled Kelvin waves in nature. It is important to investigate discrepancies of

565 flow fields between the slow and fast propagation speed scenarios. Specifically, a fast propagation
566 scenario (18 ms^{-1} for mean heating and mesoscale heating envelope) is considered here. All the
567 other model setup is the same as Sec.4.

568 Fig.10a-d shows potential temperature anomalies induced by mean heating and eddy terms.
569 Similar to Fig.6a, potential temperature anomalies induced by mean heating in Fig.10a also have an
570 upward/westward tilted vertical profile in the fast scenario. The maximum magnitude of potential
571 temperature anomalies in the fast propagation case ($1.2K$) is much stronger than that in the slow
572 propagation speed ($0.8K$). In contrast to that, potential temperature anomalies induced by eddy
573 terms in both cases have similar maximum magnitudes. The total potential temperature anomalies
574 in Fig.10c has an clear upward/westward tilt. A further investigation for potential temperature
575 anomalies in different propagation speed scenarios are shown in Fig.10d. It can be concluded
576 that mean heating tends to induce stronger potential temperature anomalies in faster propagation
577 speeds (less than 25 ms^{-1}), while anomalies induced by eddy terms do not change much. The
578 propagation speed 25 ms^{-1} as a threshold corresponds to the phase speed of gravity waves in the
579 second baroclinic modes as discussed in Sec.2.

580 Fig.10e-h shows zonal velocity induced by upward/westward mean heating and eddy terms.
581 Similar to Fig.7a in the slow scenario, zonal velocity induced by mean heating in Fig.10e also has
582 an upward/westward tilted vertical profile in the fast scenario. However, the maximum magnitude
583 of both the easterlies and westerlies in the fast scenario is 3.5 ms^{-1} , much stronger than that
584 in the slow propagation speed (2.5 ms^{-1}). The overall spatial pattern of zonal velocity induced
585 by eddy terms in Fig.10f and the total zonal velocity in Fig.10g are quite similar to that in Fig.7c
586 and Fig.7e. A further investigation for zonal velocity in different propagation speed scenarios are
587 shown in Fig.10h. Similarly, it can be concluded that mean heating tends to induce stronger zonal

588 velocity in the faster propagation speed (less than 25 ms^{-1}), while that induced by eddy terms does
589 not change much.

590 7. Westward-Moving Mesoscale Heating in An Upward/Westward Tilt

591 One crucial feature of self-similarity of tropical convection is the front-to-rear tilted verti-
592 cal structure, which is commonly observed in MCSs (Houze 2004) and CCEWs (Kiladis et al.
593 2009). In the two-dimensional cloud resolving simulations (Grabowski and Moncrieff 2001), the
594 westward-moving systems are characterized by low-level inflow and upper-level outflow in an up-
595 ward/eastward tilt. It is important to know why mesoscale systems with an upward/westward tilt
596 are unfavorable and what is the underlying physical mechanism to avoid them. In order to pre-
597 scribe such a westward-moving mesoscale system in an upward/westward tilt, here the expression
598 of mesoscale heating is the same as Eq.19 except that the sign of phase shift is reversed $\phi_0 = \frac{\pi}{4}$.
599 According to Eqs.12 and Eq.14, after switching the sign of phase shift parameter ϕ_0 from $-\frac{\pi}{4}$ to $\frac{\pi}{4}$,
600 both vertical profiles and magnitudes of EMT and EHT remains the same except for their signs. In
601 other words, the synoptic-scale circulation response to eddy terms including potential temperature
602 anomalies in Fig.6c and zonal velocity in Fig.7c has the same spatial pattern but the opposite sign.

603 Fig.11a shows relative location between mean heating and potential temperature anomalies in-
604 duced by eddy terms in the longitude-height diagram at 41.5 *day*. In the lower troposphere, posi-
605 tive potential temperature anomalies induced by eddy terms are located to the east of mean heating
606 as it propagates eastward. In a moist environment, such positive potential temperature anomalies
607 can suppress convection through decreasing convective available potential energy (CAPE), in-
608 creasing saturation rate of water vapor and convective inhibition (CIN). In the reality, the leading
609 edge of the synoptic-scale convective envelope mostly consists of shallow congestus clouds in
610 the lower troposphere. Therefore, the upscale impact of mesoscale fluctuations of momentum and

611 temperature tends to suppress shallow convection in the leading edge and further destroy the multi-
612 scale coherent structure of synoptic-scale equatorial waves. In the middle troposphere, potential
613 temperature anomalies induced by eddy terms are out of phase with the mean heating. There
614 are negative (positive) anomalies to the east (west), which favor (suppress) convection to the east
615 (west). In the upper troposphere, the negative potential temperature anomalies in the upper-level
616 mean heating region provide favorable conditions for stratiform convection.

617 Fig.11b shows relative location of potential temperature anomalies induced by mean heating
618 and eddy terms in the longitude-height diagram at 41.5 *day*. In the lower troposphere in Fig.11c,
619 potential temperature anomalies induced by eddy terms tend to strengthen those induced by mean
620 heating, resulting in strong low-level anomalies with their maximum around 0.67K. In the middle
621 troposphere, potential temperature anomalies induced by eddy terms have dominate magnitudes
622 and reverse the sign of those induced by mean heating. In connection to those in the upper tro-
623 posphere, eddy terms induce upward/eastward tilted potential temperature anomalies in the upper
624 levels. Fig.11 shows the total potential temperature anomalies induced by both mean heating and
625 eddy terms. Such spatial pattern of potential temperature anomalies are quite different from Fig.6f
626 from the study (Grabowski and Moncrieff 2001).

627 **8. Concluding Discussion**

628 Synoptic-scale equatorial waves coupled with tropical convection are typically organized in a
629 hierarchy of multiple spatial and temporal scales (Kiladis et al. 2009). In particular, convec-
630 tively coupled Kelvin waves are characterized by a eastward-moving convective envelope with
631 embedded westward-moving mesoscale disturbances (Straub and Kiladis 2002). Such multi-scale
632 coherent structure of tropical convection is simulated in the two-dimensional cloud resolving sim-
633 ulations (Grabowski and Moncrieff 2001). The first goal of this paper is using the two-dimensional

634 MESD model to capture the organized structure of large-scale circulation with westward-moving
635 mesoscale systems embedded in an eastward-moving synoptic-scale convective envelope as sim-
636 ulated in (Grabowski and Moncrieff 2001). The second goal is to assess the upscale impact of
637 mesoscale disturbances on the behavior of synoptic-scale circulation through eddy transfer of mo-
638 mentum and temperature. The third goal aims at understanding how much of synoptic-scale circu-
639 lation is induced by eddy transfer of momentum and temperature rather than synoptic-scale mean
640 heating.

641 The two-dimensional MESD model is reduced from the three-dimensional MESD model (Ma-
642 jda 2007) by following several simplifying assumptions. After ignoring all the forcing terms,
643 both mesoscale and synoptic-scale governing equations share the same gravity wave equations
644 as their dynamic core, indicating the self-similarity of flow fields and tropical convection as em-
645 phasized in (Majda 2007). In particular, the synoptic-scale equations are forced by mean heat-
646 ing and eddy transfer of momentum and temperature, where the latter represents upscale impact
647 of mesoscale disturbances on the synoptic-scale circulation. After implementing similar model
648 setup as (Grabowski and Moncrieff 2001), the synoptic-scale circulation response from the two-
649 dimensional MESD model is directly compared with results from (Grabowski and Moncrieff
650 2001). Besides, three different scenarios with elevated upright mean heating, a faster propagat-
651 ing convective envelope and westward-moving mesoscale heating in an upward/westward tilt are
652 considered and used to achieve those goals as mentioned before.

653 On the mesoscale, as driven by a prescribed westward-moving top-heavy mesoscale heating in
654 an upward/eastward tilt, the mesoscale flow fields share many similar features with the westward-
655 moving mesoscale systems in (Grabowski and Moncrieff 2001), including the tilted low-level in-
656 flow and upper-level outflow. The associated EMT is characterized by upper-level (low-level) east-
657 ward (westward) momentum forcing, while EHT is manifested by mid-level cooling and upper-

658 level/low-level heating. The relative strength of synoptic-scale circulation response induced by
659 EMT and EHT is determined by the propagation speed of mesoscale heating in the MESD model.

660 On the synoptic scale, mesoscale heating is modulated by a zonally varying large-scale envelope,
661 which propagates eastward in the same speed as the mean heating. Many features of flow fields
662 simulated by Grabowski and Moncrieff (2001) are captured here. First, the domain-averaged zonal
663 velocity at each level is characterized by upper-level westerlies and low-level easterlies, which
664 is just generated by EMT. Secondly, EMT and EHT tend to induce synoptic-scale circulation
665 in the opposite signs, although that induced by EMT dominates in the slow mesoscale heating
666 scenario. The total potential temperature anomalies induced by mean heating and eddy terms
667 have an upward/westward tilt. Also, the total zonal velocity is characterized by westerlies on top
668 of easterlies in tilted vertical structure. Thirdly, compared with eddy terms, synoptic-scale mean
669 heating induces tilted vertical motion in much stronger magnitude, which is quite different from
670 the upright vertical motion as simulated in (Grabowski and Moncrieff 2001).

671 In the scenario with elevated upright mean heating, the synoptic-scale circulation response to
672 mean heating and eddy terms includes potential temperature anomalies and zonal velocity in an
673 upward/westward tilt and upright vertical velocity, similar to that in (Grabowski and Moncrieff
674 2001). Such a scenario highlights the significant upscale impact of mesoscale disturbances on the
675 synoptic-scale circulation and supports the hypothesis that the front-to-rear tilted vertical structure
676 of large-scale flow field can be contributed by eddy terms, even in the existence of upright mean
677 heating.

678 In the scenario with a faster propagating convective envelope (less than 25 ms^{-1}) in Fig.10,
679 the synoptic-scale circulation response to mean heating becomes stronger, while that driven by
680 eddy terms does not change much. In other words, the upscale impact of mesoscale disturbances
681 becomes less important as the propagation speed of convective envelope increases. Such a dis-

682 cussion can explain discrepancies of numerical results in cloud resolving models. For example, a
683 strong easterly mean flow (-10 ms^{-1}) for the trade wind regime is assumed in the simulations by
684 Grabowski and Moncrieff (2001), where CMT significantly modify the momentum budget of the
685 large-scale flow and induces vertical shear of zonal momentum. The large-scale organization of
686 convection propagates eastward slowly (8 ms^{-1}), in agreement with the conclusion that upscale
687 impact of mesoscale disturbances plays an important role in the slow propagation scenario. In
688 contrast, zero mean flow for the state of rest regime is assumed in the simulations by Tulich and
689 Mapes (2008), where the upscale transport of horizontal momentum by coherent eddy circulations
690 is found to be small. The horizontally propagating wave packets with roughly a full-wavelength
691 structure in the troposphere has phase speed in the range $16\text{-}18 \text{ ms}^{-1}$, in agreement with the con-
692 clusion from the MESD model that the synoptic-scale circulation response to the mean heating
693 dominates in the fast propagation scenario.

694 In the scenario with upward/westward tilted mesoscale heating, positive potential temperature
695 anomalies are induced in the leading edge of the synoptic-scale mean heating by the upscale im-
696 pact of mesoscale disturbances, which tends to suppress shallow convection through increasing
697 saturation rate of vapor and CIN in a moist environment. Since shallow convection in the leading
698 edge of large-scale convective envelope serves to moisten and precondition deep convection, such
699 unfavorable conditions will destroy the multi-scale coherent structure of synoptic-scale equatorial
700 waves, explaining the fact that westward-moving mesoscale heating in an upward/westward tilt is
701 rarely observed in reality.

702 This study based on a simple multi-scale model has several implication for physical interpre-
703 tation and comprehensive numerical models. In particular, the explicit expressions for EMT and
704 EHT provide assessment of upscale impact of mesoscale disturbances on the synoptic-scale circu-
705 lation in a transparent fashion, which should be useful to improve convective parameterization of

706 organized tropical convection in the GCMs. The two-dimensional MESD model under the current
707 model setup can also be generalized in several ways. One promising research direction is to con-
708 sider the original three-dimensional MESD model with Coriolis force on the synoptic scale. In the
709 existence of Kelvin waves, Rossby waves, MRG waves and Gravity waves, more realistic features
710 are expected to be reproduced in this three-dimensional MESD model such as the zonal asymme-
711 try of flow fields between eastward and westward propagating scenarios. Besides, by considering
712 an off-equator convective envelope, it is also possible to mimic the scenario when CCEWs goes
713 along the climatological ITCZ over the central and eastern Pacific.

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715 search ONR MURI N00014-12-1-0912, and Q.Y. is supported as a graduate research assistant on
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717 APPENDIX

718 Gravity waves on the mesoscale and their upscale fluxes

719 It is proved here that free gravity waves in Eqs.1a-1d cannot generate nonzero EMT and EHT. In
720 order to solve this system analytically, an ansatz for plane waves in the baroclinic mode is assumed
721 as follows,

$$f = \tilde{f} e^{i(kx - \omega\tau)} \cos(qz), f \in \{u, v, p\} \quad (\text{A1})$$

$$g = \tilde{g} e^{i(kx - \omega\tau)} [-q \sin(qz)], g \in \{w, \theta\} \quad (\text{A2})$$

722 where k is the wavenumber and ω is frequency. $q = 1, 2, 3, \dots$ is the vertical index.

723 As shown in Eq.12 and Eq.14, both EMT and EHT are in the quadratic form between vertical
724 velocity w and zonal velocity u (or potential temperature anomalies θ). According to the definition
725 of mesoscale zonal and temporal averaging in Eqs.3-4, two necessary conditions for nonzero eddy

726 terms are,

$$|k_1| = |k_2| \quad (\text{A3})$$

$$|\omega_1| = |\omega_2| \quad (\text{A4})$$

727 where k_1, ω_1 are for vertical velocity w and k_2, ω_2 are for u (or θ). According to the dispersion
728 relation in Eq.9, the necessary conditions in Eqs.A3-A4 further imply,

$$|q_1| = |q_2| \quad (\text{A5})$$

729 which requires that w and u (or θ) are in the same baroclinic mode.

730 After plugging the ansatz in Eqs.A1-A2, the mesoscale equations in Eqs.1a-1d can be reduced
731 to,

$$-i\omega\tilde{u} + ik\tilde{p} = 0 \quad (\text{A6a})$$

$$-i\omega\tilde{\theta} + \tilde{w} = 0 \quad (\text{A6b})$$

$$\tilde{p} = \tilde{\theta} \quad (\text{A6c})$$

$$ik\tilde{u} - q^2\tilde{w} = 0 \quad (\text{A6d})$$

732 According to Eq.A6b and Eq.A6d, it can be concluded that w and u (w and θ) are out of phase by
733 $\frac{\pi}{2}$. Such an out-of-phase relation means that the value of EMT and EHT after taking mesoscale
734 averaging should vanish.

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829 **LIST OF TABLES**

830 **Table 1.** The parameters, constant and scaling of physical variables in the two-
831 dimensional MESD model. 43

832 **Table 2.** Grid number and time steps in solving the synoptic-scale equations in the two-
833 dimensional MESD model. 44

	Physical variables	Symbolic notation	Value
Synoptic scale	Zonal scale	X	1500 <i>km</i>
	Temporal scale	t	8.3 <i>h</i>
	Zonal velocity	U	5 ms^{-1}
	Vertical velocity	W	$1.6 \times 10^{-2} ms^{-1}$
	Pressure perturbation	P	$250 m^2 s^{-2}$
	Potential temperature anomalies	Θ	3.3 <i>K</i>
	Zonal momentum forcing	S^u	$15 ms^{-1} day^{-1}$
	Thermal forcing	S^θ	10 $K day^{-1}$
Mesoscale	Zonal scale	x	150 <i>km</i>
	Temporal scale	τ	50 <i>min</i>
	Zonal velocity	u	5 ms^{-1}
	Vertical velocity	w	$1.6 \times 10^{-1} ms^{-1}$
	Pressure perturbation	p	$250 m^2 s^{-2}$
	Potential temperature anomalies	θ	3.3 <i>K</i>
	Zonal momentum forcing	s_u	$150 ms^{-1} day^{-1}$
	Thermal forcing	s_θ	100 $K day^{-1}$

TABLE 1. The parameters, constant and scaling of physical variables in the two-dimensional MESD model.

Resolution	Notation	Value
Zonal length	L	20,000 <i>km</i>
Vertical length	H	15.7 <i>km</i>
Total time	T	41.5 <i>day</i>
x-grid number	N_x	201
x-grid spacing	Δx	100 <i>km</i>
z-grid number	N_z	31
z-grid spacing	Δz	0.5 <i>km</i>
Time step	N_t	3600
Time interval	Δt	16.6 <i>min</i>

834 TABLE 2. Grid number and time steps in solving the synoptic-scale equations in the two-dimensional MESD
835 model.

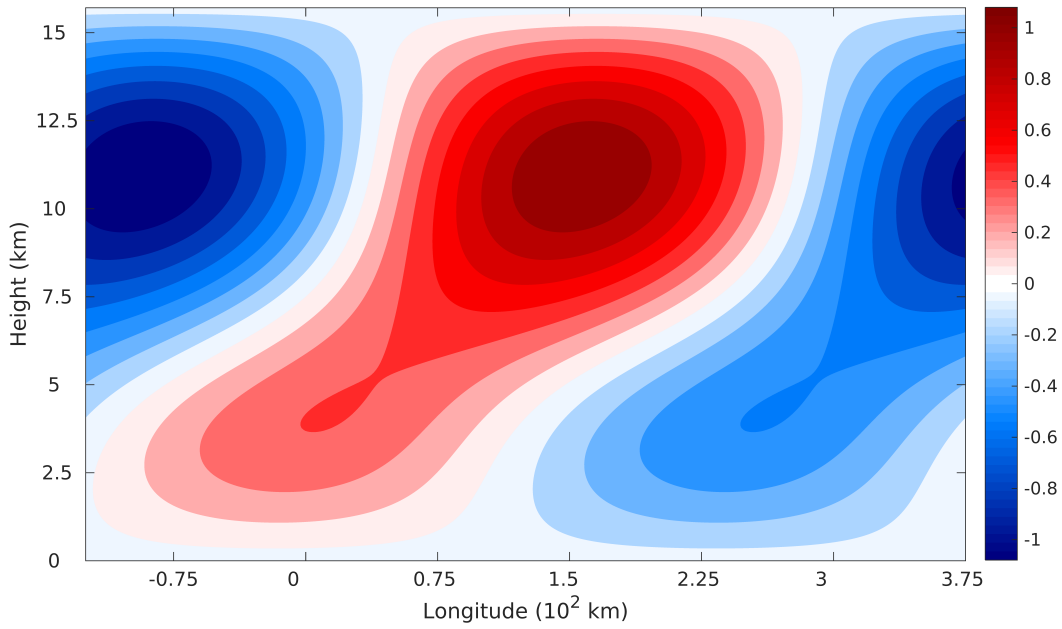
LIST OF FIGURES

836		
837	Fig. 1.	Spatial pattern of westward-moving mesoscale heating in the longitude-height diagram. One dimensionless unit corresponds to $100Kday^{-1}$ 47
838		
839	Fig. 2.	Zonal velocity, vertical velocity and potential temperature anomalies on the mesoscale in the longitude-height diagram. The panels show (a-b) zonal velocity, (c-d) streamfunction, (e) potential temperature anomalies, (f) vertical velocity. Panels (b,d) show Figure 4b and Figure 4c from the paper (Grabowski and Moncrieff 2001). The contour interval of zonal velocity is $0.98 ms^{-1}$. The dimensional units of zonal velocity, vertical velocity and potential temperature anomalies are ms^{-1} , ms^{-1} , K 48
840		
841		
842		
843		
844		
845	Fig. 3.	Vertical structure of (a) eddy momentum transfer and (b) eddy heat transfer. In each panel, the red curve is for eddy fluxes and blue curve is for vertical gradient of eddy fluxes in a minus sign. One dimensionless unit of eddy momentum transfer is $15ms^{-1}day^{-1}$ and that of eddy heat transfer is $10Kday^{-1}$ 49
846		
847		
848		
849	Fig. 4.	Spatial distribution of momentum flux and zonal velocity at 41.5 day and time series of domain-averaged zonal momentum. The left column shows numerical solutions and the right column shows Figure 16a and Figure 17a from the paper (Grabowski and Moncrieff 2001). The panels show (a-b) momentum flux, (c) zonal velocity, (d-e) domain-averaged zonal momentum. The contour interval in panels is (a) $0.03 m^2s^{-2}$, (b) $0.02 Nm^{-2}$, (d) $0.42 ms^{-1}$, (e) $0.5 ms^{-1}$. The dimensional units of momentum flux, zonal velocity in the numerical solutions are m^2s^{-2} and ms^{-1} 50
850		
851		
852		
853		
854		
855		
856	Fig. 5.	Mean heating (color) and zonal/vertical velocity (arrow) in the longitude-height diagram at 41.5 day. Only the solutions in the longitude range from $6.87 \times 10^3 km$ to $12.84 \times 10^3 km$ are plotted here. The maximum magnitude of zonal and vertical velocity are $1.92ms^{-1}$ and $0.89cms^{-1}$. The dimensional unit of mean heating is $10 Kday^{-1}$ 51
857		
858		
859		
860	Fig. 6.	Potential temperature anomalies (from the domain average at each level) in the longitude-height diagram at 41.5 day. The panel (f) shows Figure 13a from the paper (Grabowski and Moncrieff 2001). The rest panels show potential temperature anomalies induced by (a) mean heating, (b) eddy momentum transfer, (c) eddy momentum transfer and eddy heat transfer (contour), (d) eddy heat transfer, (e) total. The contour interval in panel (c) $0.05 K$ and that in panel (f) is $0.2 K$. The dimensional unit is K 52
861		
862		
863		
864		
865		
866	Fig. 7.	Zonal velocity in the longitude-height diagram at 41.5 day. The panel (f) shows Figure 14a from the paper (Grabowski and Moncrieff 2001). The rest panels show zonal velocity induced by (a) mean heating, (b) eddy momentum transfer, (c) eddy momentum transfer and eddy heat transfer (contour), (d) eddy heat transfer, (e) total. The contour interval in panel (c) is $0.19 ms^{-1}$ and that in panel (f) is $1 ms^{-1}$. The dimensional unit is ms^{-1} 53
867		
868		
869		
870		
871	Fig. 8.	Vertical velocity in the longitude-height diagram at 41.5 day. The panel (d) shows Figure 14b from the paper (Grabowski and Moncrieff 2001). The rest panels show vertical velocity induced by (a) eddy momentum transfer, (b) eddy heat transfer, (c) mean heating. The dimensional unit is $10^{-2}ms^{-1}$. The vertical velocity induced by mean heating in panel (c) has a upward/westward tilt in dominant magnitude, which is different from panel (d). A similar scenario with upright mean heating is discussed in Sec.5. 54
872		
873		
874		
875		
876		
877	Fig. 9.	Upright mean heating, potential temperature anomalies (from the domain average at each level), zonal and vertical velocity in the longitude-height diagram at 41.5 day. The panels show (a) upright mean heating, (b) vertical velocity, (c) potential temperature anomalies
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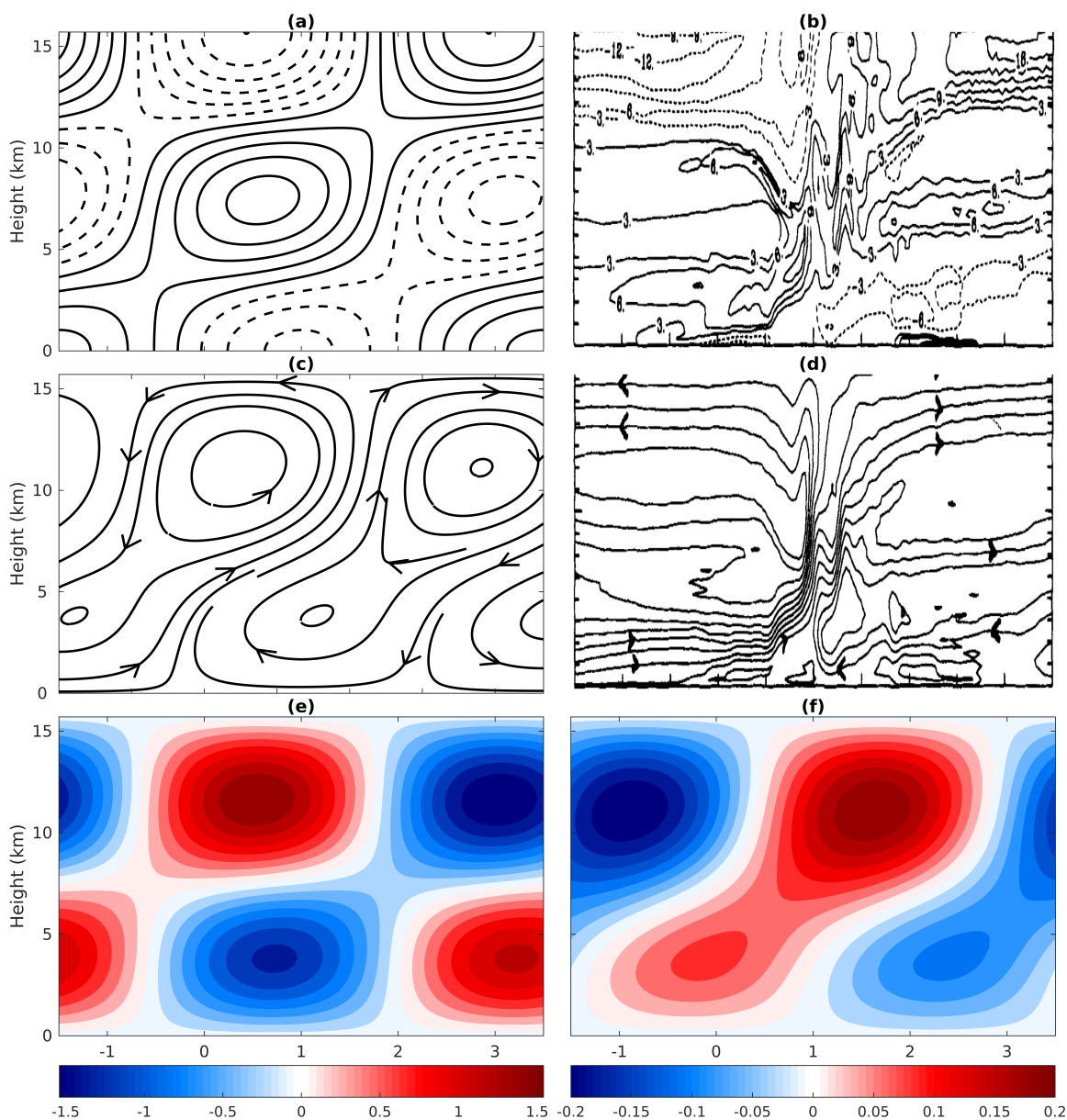
880 induced by eddy terms (contour) and mean heating (color), (d) total potential temperature
 881 anomalies. (e) zonal velocity induced by eddy terms (contour) and mean heating (color),
 882 (f) total zonal velocity. The dimensional unit of mean heating, vertical velocity, potential
 883 temperature anomalies and zonal velocity are 10 Kday^{-1} , 10^{-2} ms^{-1} , K , ms^{-1} respectively. 55

884 **Fig. 10.** Potential temperature anomalies (from the domain average at each level) and zonal velocity
 885 in the longitude-height diagram at 41.5 *day*. The left panels show potential temperature
 886 anomalies induced by (a) mean heating, (b) eddy terms, (c) total. The right panels (d-f)
 887 are similar to (a-c) but for zonal velocity. Panel (d) shows potential temperature anomalies
 888 in Frobenius norm in different propagation speeds and panel (h) shows the same but for
 889 zonal velocity. Panels (d) and (h) share the same legend and the two dashed lines denote
 890 the phase speeds of gravity waves in the second (25 ms^{-1}) and third (16.7 ms^{-1}) baroclinic
 891 modes. The dimensional units of potential temperature anomalies and zonal velocity are K
 892 and ms^{-1} 56

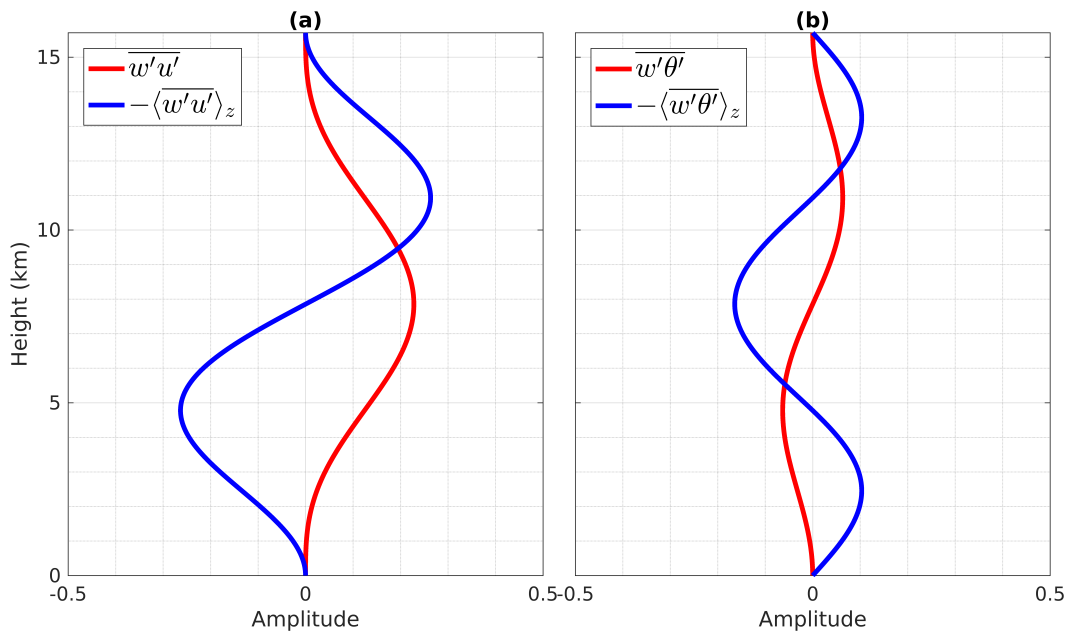
893 **Fig. 11.** Potential temperature anomalies (from the domain average at each level) in the longitude-
 894 height diagram at 41.5 *day*. Panel (a) shows potential temperature anomalies induced by
 895 eddy terms (contour), and mean heating (color). The right panels show potential temperature
 896 anomalies induce by (b) mean heating (color) and eddy terms (contour), (c) total. The
 897 dimensional units of mean heating and potential temperature anomalies are 10 Kday^{-1} , K
 898 respectively. 57



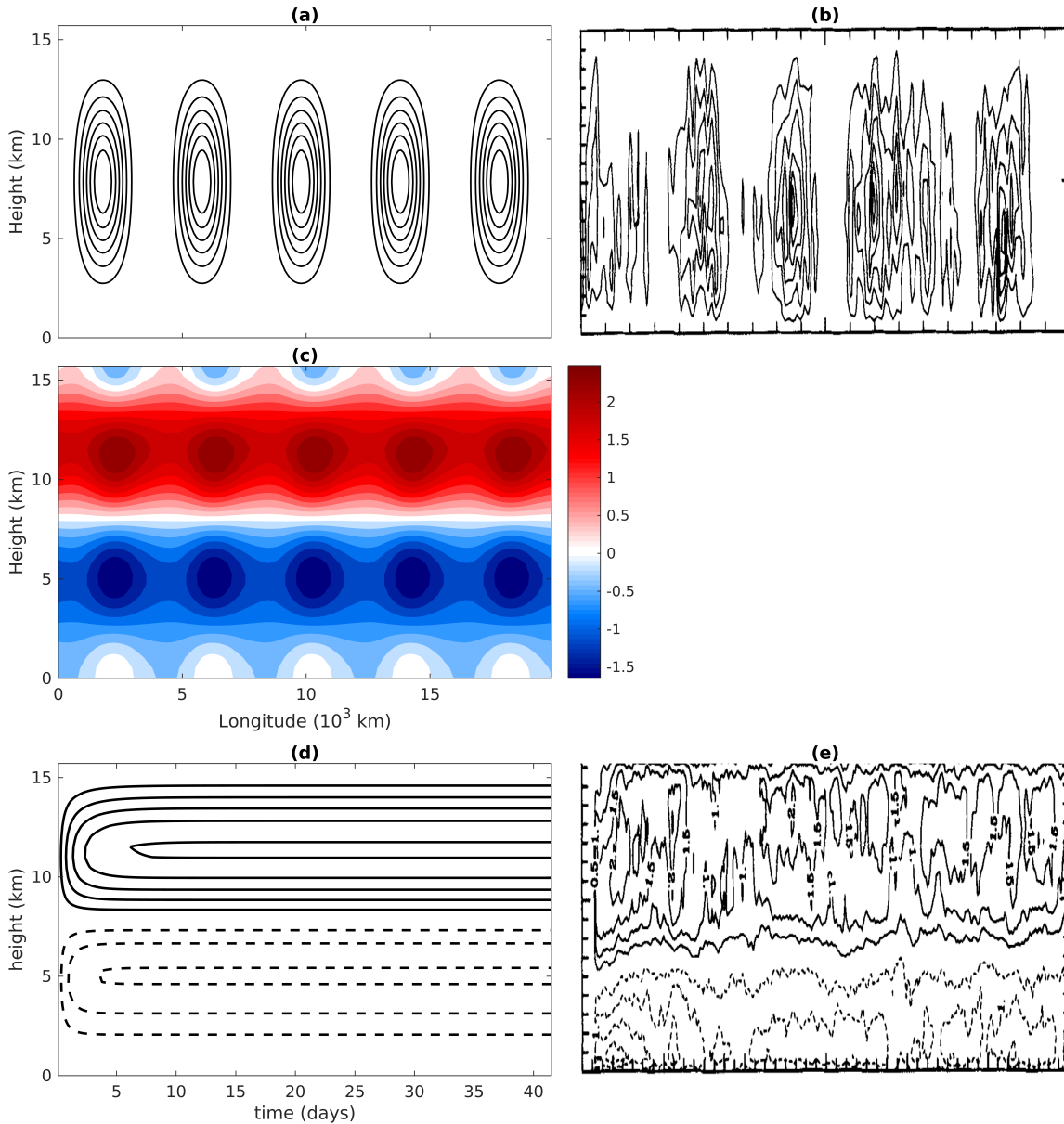
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900 sionless unit corresponds to $100Kday^{-1}$.



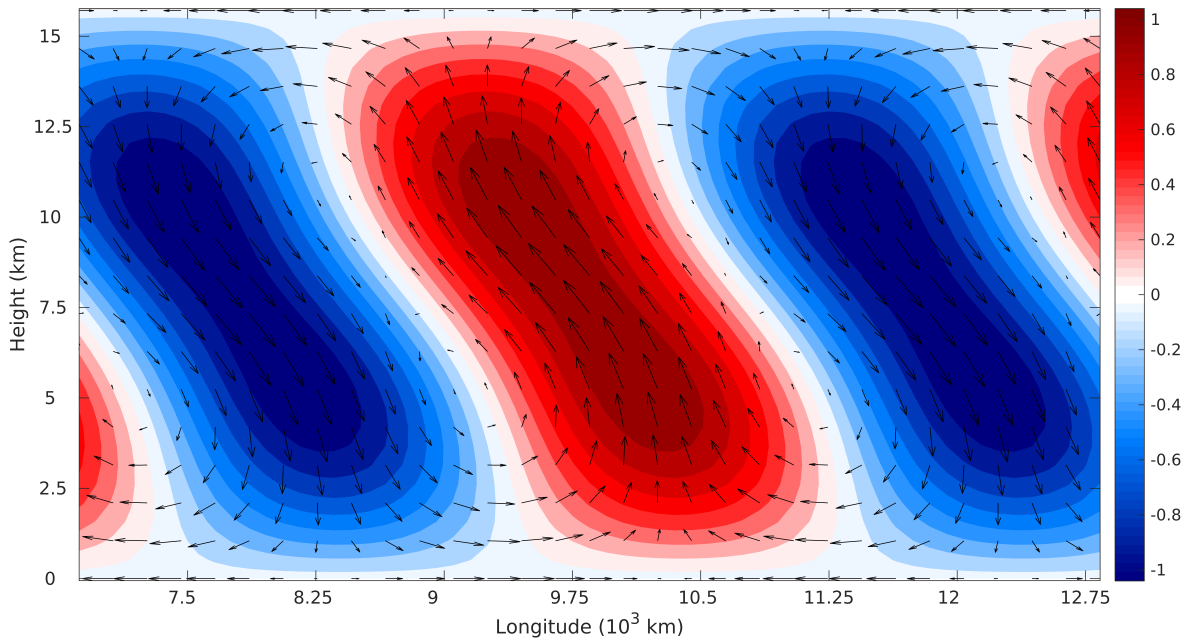
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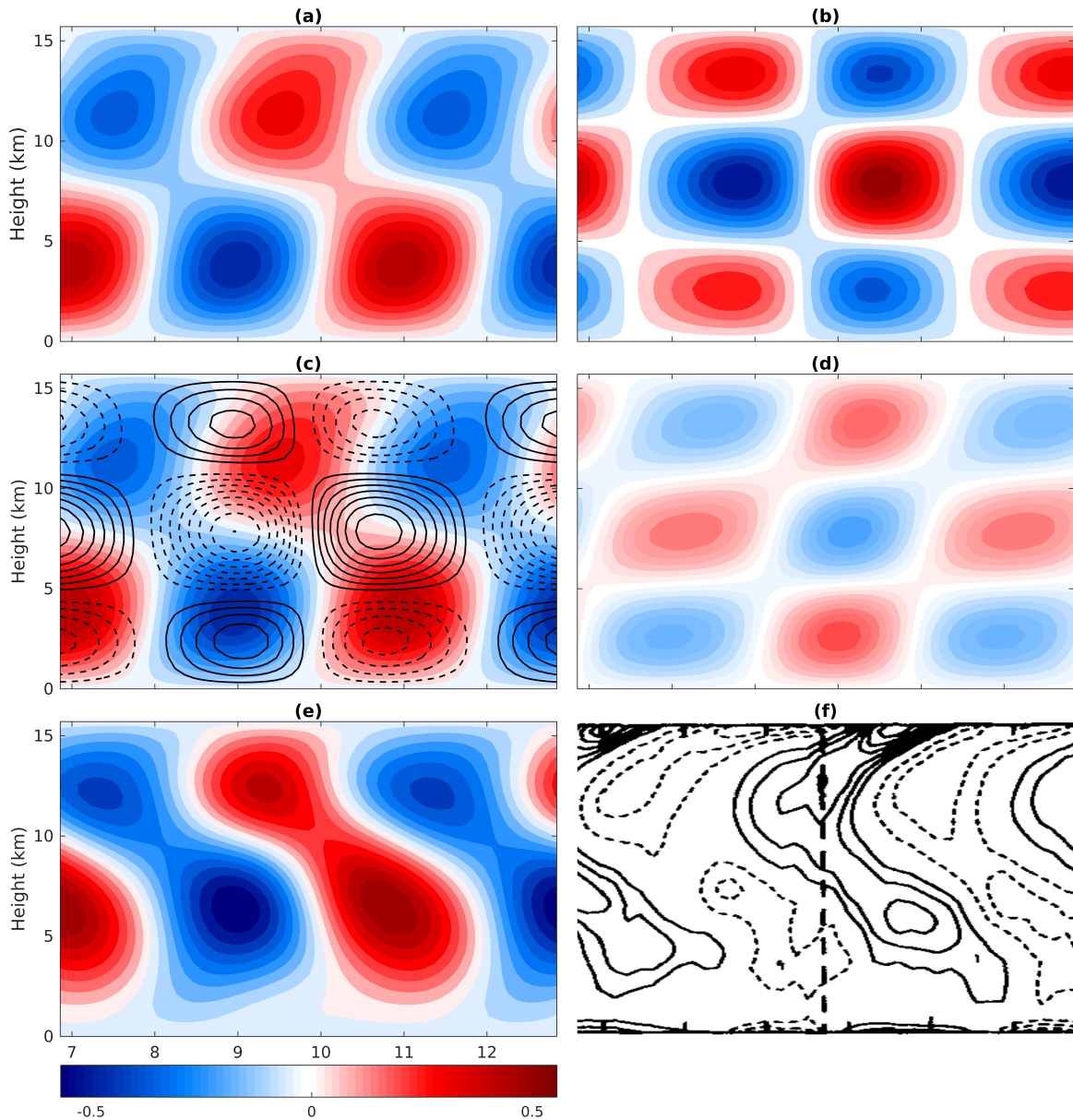
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 907 curve is for eddy fluxes and blue curve is for vertical gradient of eddy fluxes in a minus sign. One dimensionless
 908 unit of eddy momentum transfer is $15ms^{-1}day^{-1}$ and that of eddy heat transfer is $10Kday^{-1}$.



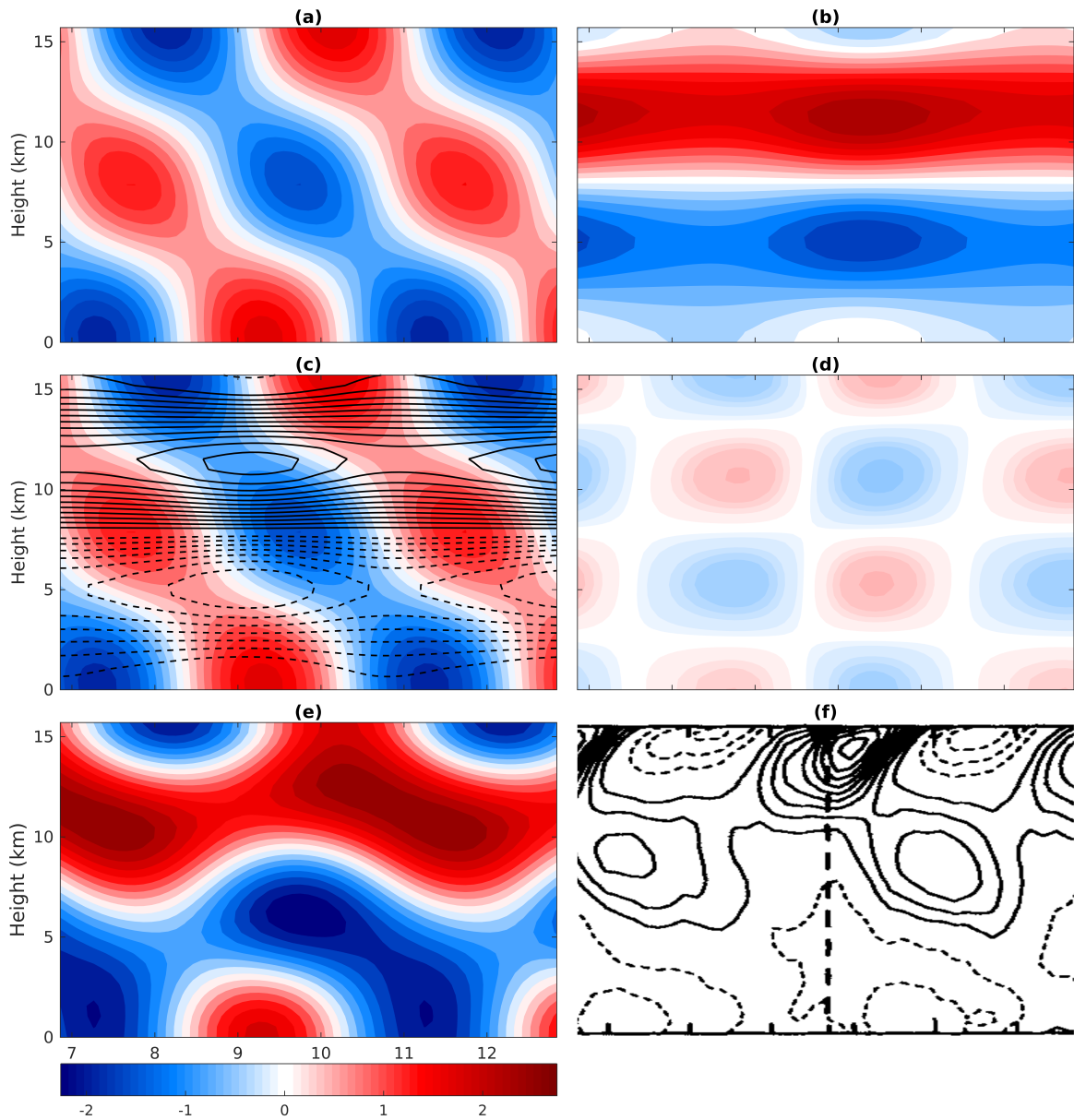
909 FIG. 4. Spatial distribution of momentum flux and zonal velocity at 41.5 day and time series of domain-
 910 averaged zonal momentum. The left column shows numerical solutions and the right column shows Figure
 911 16a and Figure 17a from the paper (Grabowski and Moncrieff 2001). The panels show (a-b) momentum flux,
 912 (c) zonal velocity, (d-e) domain-averaged zonal momentum. The contour interval in panels is (a) $0.03 \text{ m}^2\text{s}^{-2}$,
 913 (b) 0.02 Nm^{-2} , (d) 0.42 ms^{-1} , (e) 0.5 ms^{-1} . The dimensional units of momentum flux, zonal velocity in the
 914 numerical solutions are m^2s^{-2} and ms^{-1} .



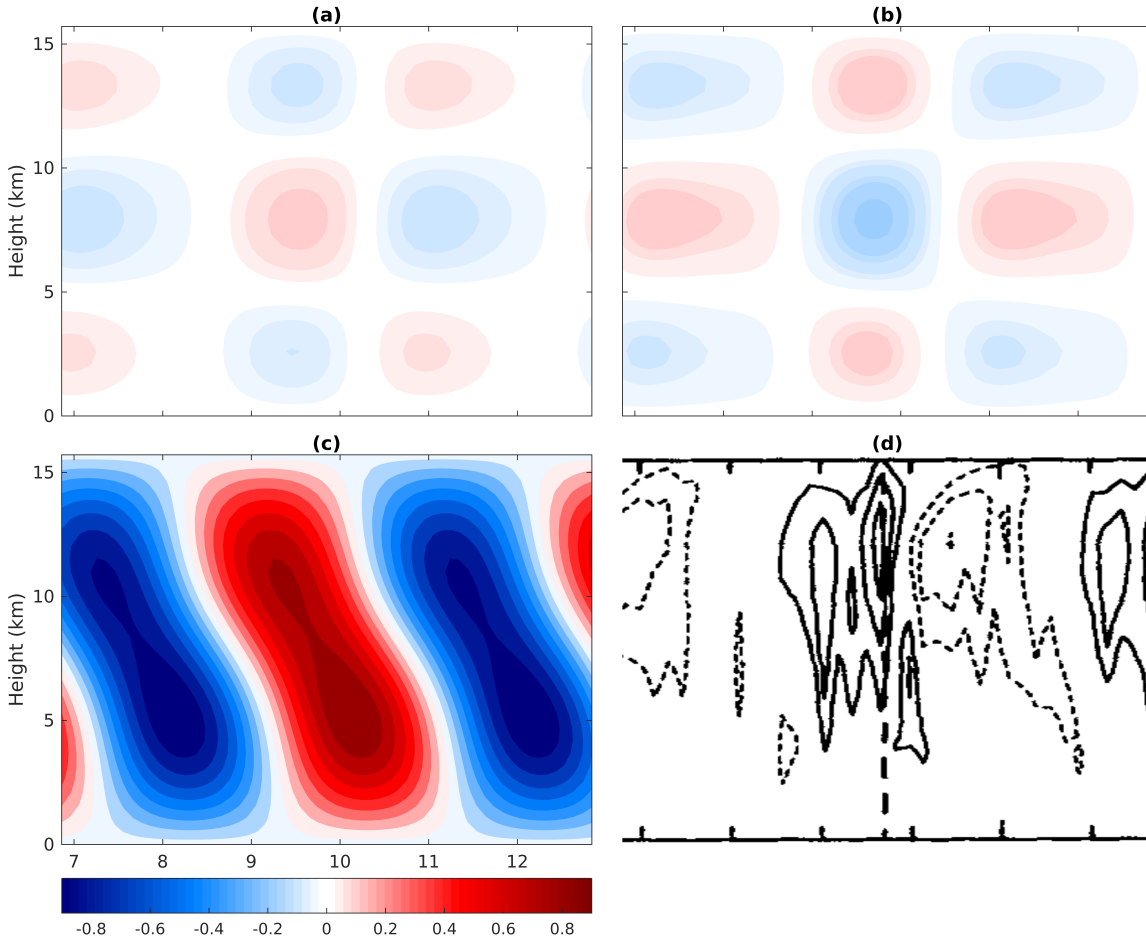
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 917 magnitude of zonal and vertical velocity are $1.92ms^{-1}$ and $0.89cms^{-1}$. The dimensional unit of mean heating is
 918 $10 Kday^{-1}$.



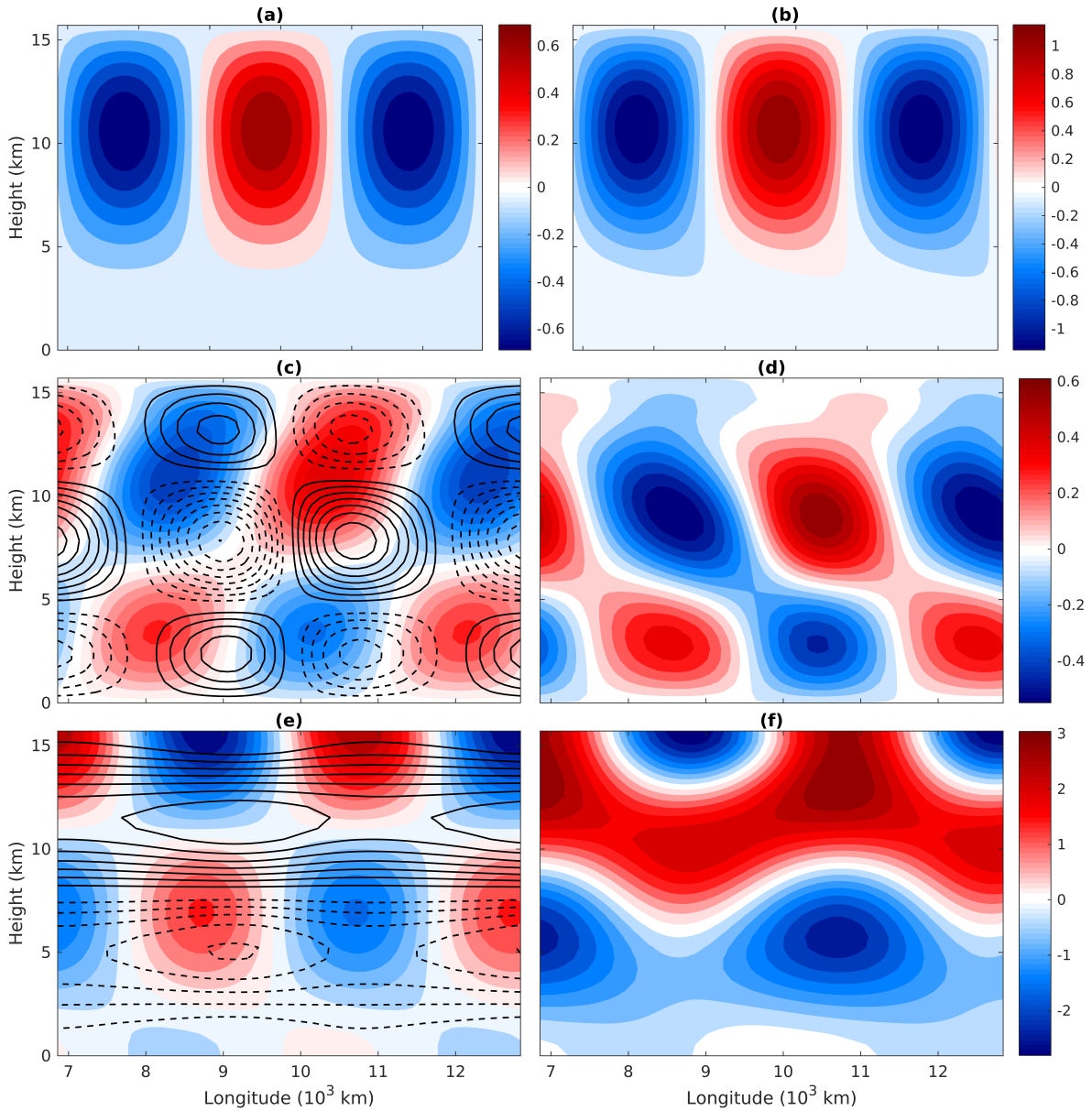
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 920 diagram at 41.5 *day*. The panel (f) shows Figure 13a from the paper (Grabowski and Moncrieff 2001). The
 921 rest panels show potential temperature anomalies induced by (a) mean heating, (b) eddy momentum transfer, (c)
 922 eddy momentum transfer and eddy heat transfer (contour), (d) eddy heat transfer, (e) total. The contour interval
 923 in panel (c) 0.05 *K* and that in panel (f) is 0.2 *K*. The dimensional unit is *K*.



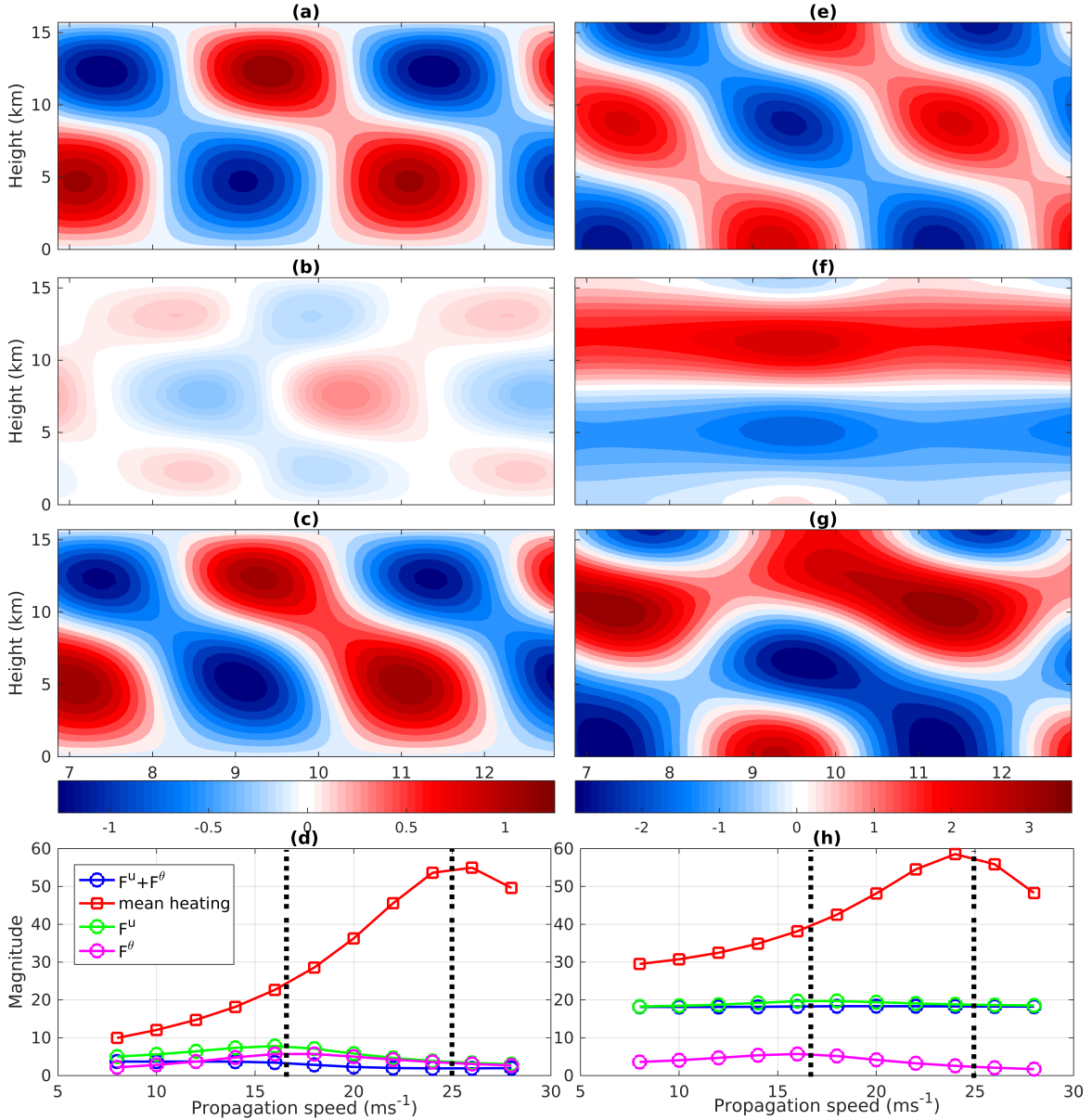
924 FIG. 7. Zonal velocity in the longitude-height diagram at 41.5 *day*. The panel (f) shows Figure 14a from the
 925 paper (Grabowski and Moncrieff 2001). The rest panels show zonal velocity induced by (a) mean heating, (b)
 926 eddy momentum transfer, (c) eddy momentum transfer and eddy heat transfer (contour), (d) eddy heat transfer,
 927 (e) total. The contour interval in panel (c) is 0.19 ms^{-1} and that in panel (f) is 1 ms^{-1} . The dimensional unit is
 928 ms^{-1} .



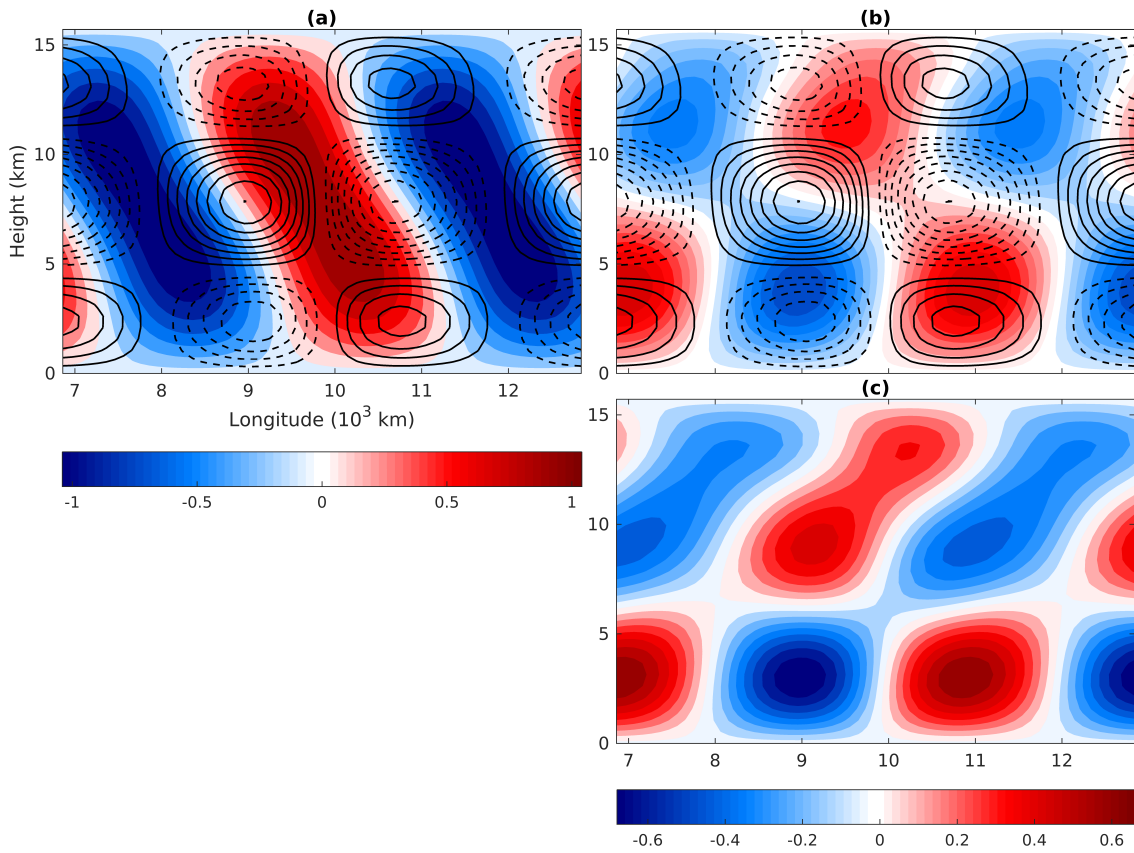
929 FIG. 8. Vertical velocity in the longitude-height diagram at 41.5 *day*. The panel (d) shows Figure 14b from the
 930 paper (Grabowski and Moncrieff 2001). The rest panels show vertical velocity induced by (a) eddy momentum
 931 transfer, (b) eddy heat transfer, (c) mean heating. The dimensional unit is 10^{-2}ms^{-1} . The vertical velocity
 932 induced by mean heating in panel (c) has a upward/westward tilt in dominant magnitude, which is different from
 933 panel (d). A similar scenario with upright mean heating is discussed in Sec.5.



934 FIG. 9. Upright mean heating, potential temperature anomalies (from the domain average at each level), zonal
 935 and vertical velocity in the longitude-height diagram at 41.5 *day*. The panels show (a) upright mean heating, (b)
 936 vertical velocity, (c) potential temperature anomalies induced by eddy terms (contour) and mean heating (color),
 937 (d) total potential temperature anomalies. (e) zonal velocity induced by eddy terms (contour) and mean heating
 938 (color), (f) total zonal velocity. The dimensional unit of mean heating, vertical velocity, potential temperature
 939 anomalies and zonal velocity are 10 K day^{-1} , 10^{-2} ms^{-1} , K , ms^{-1} respectively.



940 FIG. 10. Potential temperature anomalies (from the domain average at each level) and zonal velocity in the
 941 longitude-height diagram at 41.5 day. The left panels show potential temperature anomalies induced by (a)
 942 mean heating, (b) eddy terms, (c) total. The right panels (d-f) are similar to (a-c) but for zonal velocity. Panel (d)
 943 shows potential temperature anomalies in Frobenius norm in different propagation speeds and panel (h) shows
 944 the same but for zonal velocity. Panels (d) and (h) share the same legend and the two dashed lines denote the
 945 phase speeds of gravity waves in the second ($25 ms^{-1}$) and third ($16.7 ms^{-1}$) baroclinic modes. The dimensional
 946 units of potential temperature anomalies and zonal velocity are K and ms^{-1} .



947 FIG. 11. Potential temperature anomalies (from the domain average at each level) in the longitude-height
 948 diagram at 41.5 *day*. Panel (a) shows potential temperature anomalies induced by eddy terms (contour), and
 949 mean heating (color). The right panels show potential temperature anomalies induced by (b) mean heating (color)
 950 and eddy terms (contour), (c) total. The dimensional units of mean heating and potential temperature anomalies
 951 are 10 K day^{-1} , K respectively.