

Coupling a Fluctuating Fluid with Suspended Structures

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Outline

- 1 Introduction
- 2 Fluctuating Hydrodynamics
- 3 Incompressible Inertial Coupling
- 4 Numerics
- 5 Outlook

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- An important feature of small-scale flows, not discussed here, is **surface/boundary effects** (e.g., slip in the contact line problem).
- Essential distinguishing feature from “ordinary” CFD: **thermal fluctuations!**
- I hope to demonstrate the general conclusion that **fluctuations should be taken into account at all level.**

Levels of Coarse-Graining

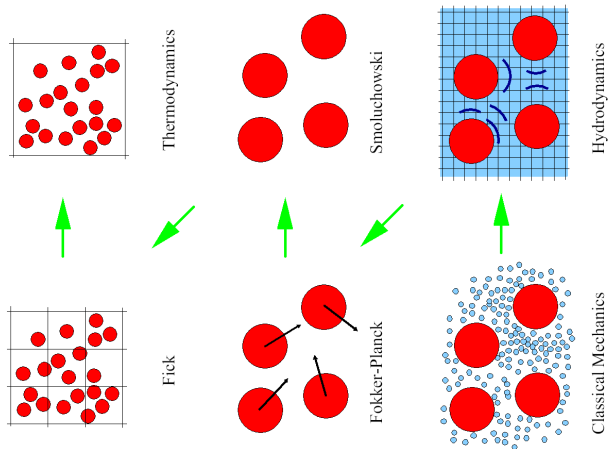
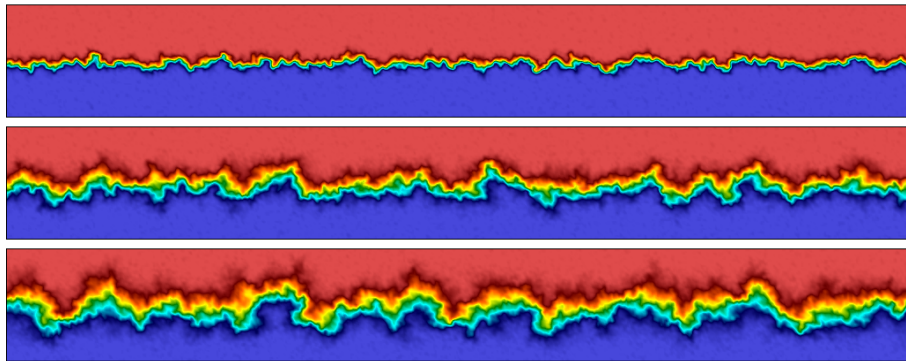


Figure: From Pep Español, “Statistical Mechanics of Coarse-Graining”

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Thermal Fluctuations Matter



Snapshots of concentration in a miscible mixture showing the development of a *rough* diffusive interface between two miscible fluids in zero gravity [1, 2, 3]. A similar pattern is seen over a broad range of Schmidt numbers and is affected strongly by nonzero gravity.

Fluctuating Navier-Stokes Equations

- We will consider a binary fluid mixture with mass **concentration** $c = \rho_1/\rho$ for two fluids that are dynamically **identical**, where $\rho = \rho_1 + \rho_2$ (e.g., **fluorescently-labeled** molecules).

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$$\begin{aligned}\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v} &= -\nabla \pi + \nu \nabla^2 \mathbf{v} + \nabla \cdot \left(\sqrt{2\nu\rho^{-1} k_B T} \mathcal{W} \right) \\ \partial_t c + \mathbf{v} \cdot \nabla c &= \chi \nabla^2 c + \nabla \cdot \left(\sqrt{2m\chi\rho^{-1} c(1-c)} \mathcal{W}^{(c)} \right),\end{aligned}$$

where the **kinematic viscosity** $\nu = \eta/\rho$, and π is determined from incompressibility, $\nabla \cdot \mathbf{v} = 0$.

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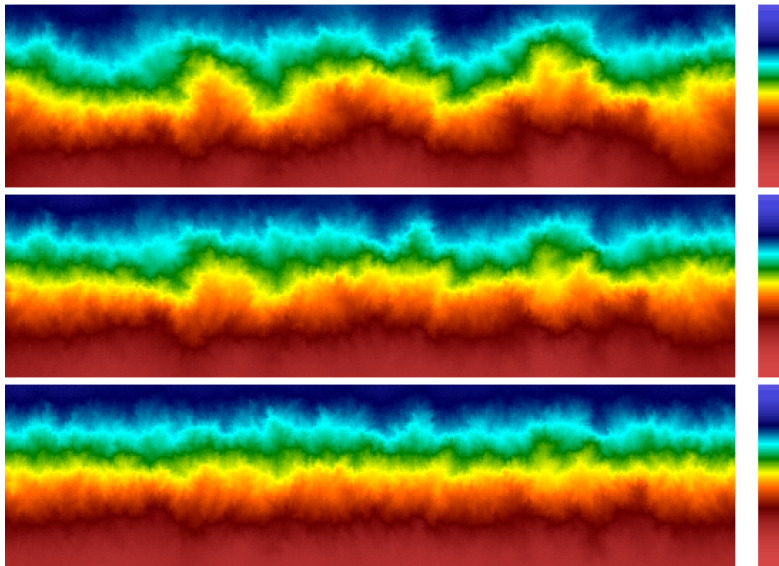
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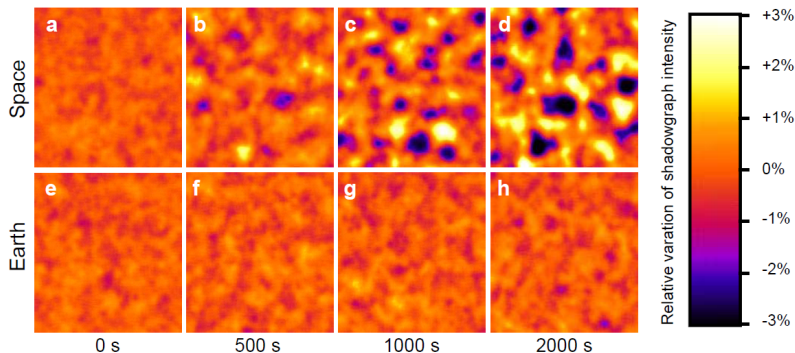
- We assume that \mathcal{W} can be modeled as spatio-temporal **white noise** (a delta-correlated Gaussian random field), e.g.,

$$\langle \mathcal{W}_{ij}(\mathbf{r}, t) \mathcal{W}_{kl}^*(\mathbf{r}', t') \rangle = (\delta_{ik} \delta_{jl} + \delta_{il} \delta_{jk}) \delta(t - t') \delta(\mathbf{r} - \mathbf{r}').$$

Fractal Fronts in Diffusive Mixing



Giant Fluctuations in Experiments



Experimental results by A. Vailati *et al.* from a microgravity environment [2] showing the enhancement of concentration fluctuations in space (box scale is **macroscopic**: 5mm on the side, 1mm thick)..

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- We have algorithms and codes to solve the compressible equations (**collocated** and **staggered grid**), and recently also the incompressible and **low Mach number** ones (staggered grid) [4, 3].
- Solving these sort of equations numerically requires paying attention to **discrete fluctuation-dissipation balance**, in addition to the usual deterministic difficulties [4].

Finite-Volume Schemes

$$c_t = -\mathbf{v} \cdot \nabla c + \chi \nabla^2 c + \nabla \cdot \left(\sqrt{2\chi} \mathbf{w} \right) = \nabla \cdot \left[-c\mathbf{v} + \chi \nabla c + \sqrt{2\chi} \mathbf{w} \right]$$

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- Generic **finite-volume spatial discretization**

$$\mathbf{c}_t = \mathbf{D} \left[(-\mathbf{V}\mathbf{c} + \mathbf{G}\mathbf{c}) + \sqrt{2\chi / (\Delta t \Delta V)} \mathbf{W} \right],$$

where $\mathbf{D} : \text{faces} \rightarrow \text{cells}$ is a **conservative** discrete divergence,
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- Advection should be **skew-adjoint** (non-dissipative) if $\nabla \cdot \mathbf{v} = 0$,

$$(\mathbf{D}\mathbf{V})^* = -(\mathbf{D}\mathbf{V}) \text{ if } (\mathbf{D}\mathbf{V}) \mathbf{1} = \mathbf{0}.$$

Weak Accuracy

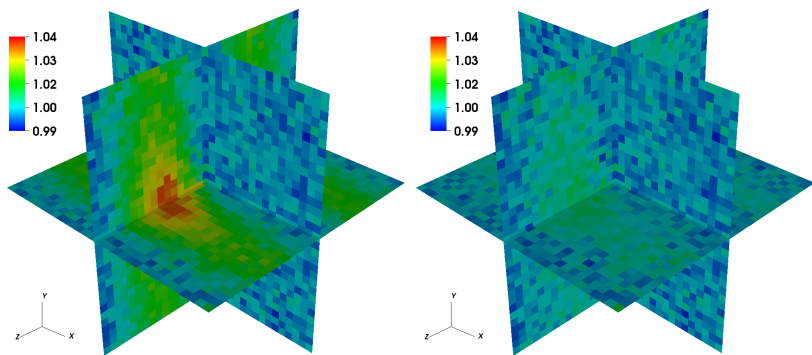


Figure: Spectral power of the first solenoidal mode for an incompressible fluid as a function of the wavenumber. The left panel is for a (normalized) time step $\alpha = 0.5$, and the right for $\alpha = 0.25$.

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- The above two conditions are **questionable at nanoscales**, but even worse, they are very hard to implement numerically in an efficient and stable manner.
- We saw already that **fluctuations should be taken into account at the continuum level**.

Brownian Particle Model

- Consider a **Brownian “particle”** of size a with position $\mathbf{q}(t)$ and velocity $\mathbf{u} = \dot{\mathbf{q}}$, and the velocity field for the fluid is $\mathbf{v}(\mathbf{r}, t)$.

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- Often presented as an interpolation function for point Lagrangian particles but here a is a **physical size** of the particle.
- We will call our particles “**blobs**” since they are not really point particles.

Local Averaging and Spreading Operators

- Postulate a **no-slip condition** between the particle and local fluid velocities,

$$\dot{\mathbf{q}} = \mathbf{u} = [\mathbf{J}(\mathbf{q})] \mathbf{v} = \int \delta_a(\mathbf{q} - \mathbf{r}) \mathbf{v}(\mathbf{r}, t) d\mathbf{r},$$

where the *local averaging* linear operator $\mathbf{J}(\mathbf{q})$ averages the fluid velocity inside the particle to estimate a local fluid velocity.

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- The **induced force density** in the fluid because of the particle is:

$$\mathbf{f} = -\lambda \delta_a(\mathbf{q} - \mathbf{r}) = -[\mathbf{S}(\mathbf{q})] \lambda,$$

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- The physical **volume** of the particle ΔV is related to the shape and width of the kernel function via

$$\Delta V = (\mathbf{JS})^{-1} = \left[\int \delta_a^2(\mathbf{r}) d\mathbf{r} \right]^{-1}. \quad (1)$$

Fluid-Structure Direct Coupling

- The equations of motion in our coupling approach are **postulated** [5] to be

$$\begin{aligned}\rho(\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v}) &= -\nabla \pi - \nabla \cdot \boldsymbol{\sigma} - [\mathbf{S}(\mathbf{q})] \boldsymbol{\lambda} + \text{thermal drift} \\ m_e \dot{\mathbf{u}} &= \mathbf{F}(\mathbf{q}) + \boldsymbol{\lambda} \\ \text{s.t. } \mathbf{u} &= [\mathbf{J}(\mathbf{q})] \mathbf{v} \text{ and } \nabla \cdot \mathbf{v} = 0,\end{aligned}$$

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- In the existing (stochastic) IBM approaches [6] **inertial effects** are ignored, $m_e = 0$ and thus $\boldsymbol{\lambda} = -\mathbf{F}$.

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- Adding the fluid and particle equations gives a **local momentum conservation law**

$$\partial_t \mathbf{p} = -\nabla \pi - \nabla \cdot \boldsymbol{\sigma} - \nabla \cdot [\rho \mathbf{v} \mathbf{v}^T + m_e \mathbf{S} (\mathbf{u} \mathbf{u}^T)] + \mathbf{S} \mathbf{F}.$$

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- For the fluid we get the effective equation

$$\rho_{\text{eff}} \partial_t \mathbf{v} = - \left[\rho (\mathbf{v} \cdot \nabla) + m_e \mathbf{S} \left(\mathbf{u} \cdot \frac{\partial}{\partial \mathbf{q}} \mathbf{J} \right) \right] \mathbf{v} - \nabla \pi - \nabla \cdot \boldsymbol{\sigma} + \mathbf{S}\mathbf{F}$$

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where the **effective mass** $m = m_e + m_f$ includes the mass of the “excluded” fluid

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- For the fluid we get the effective equation

$$\rho_{\text{eff}} \partial_t \mathbf{v} = - \left[\rho (\mathbf{v} \cdot \nabla) + m_e \mathbf{S} \left(\mathbf{u} \cdot \frac{\partial}{\partial \mathbf{q}} \mathbf{J} \right) \right] \mathbf{v} - \nabla \pi - \nabla \cdot \boldsymbol{\sigma} + \mathbf{S}\mathbf{F}$$

where the effective **mass density matrix** (operator) is

$$\rho_{\text{eff}} = \rho + m_e \mathcal{P} \mathbf{S} \mathbf{J} \mathcal{P},$$

where \mathcal{P} is the L_2 **projection operator** onto the linear subspace $\nabla \cdot \mathbf{v} = 0$, with the appropriate BCs.

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- No entropic contribution to the coarse-grained free energy because our formulation is isothermal and the particles do not have internal structure.

contd.

- A key ingredient of fluctuation-dissipation balance is that that the fluid-particle **coupling is non-dissipative**, i.e., in the absence of viscous dissipation the kinetic energy H is conserved.

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- Crucial for **energy conservation** is that $\mathbf{J}(\mathbf{q})$ and $\mathbf{S}(\mathbf{q})$ are **adjoint**, $\mathbf{S} = \mathbf{J}^*$,

$$(\mathbf{J}\mathbf{v}) \cdot \mathbf{u} = \int \mathbf{v} \cdot (\mathbf{S}\mathbf{u}) d\mathbf{r} = \int \delta_a(\mathbf{q} - \mathbf{r}) (\mathbf{v} \cdot \mathbf{u}) d\mathbf{r}. \quad (2)$$

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- The dynamics is **not incompressible in phase space** and “**thermal drift**” correction terms need to be included [6], but they turn out to **vanish** for incompressible flow (gradient of scalar).
- The spatial discretization should preserve these properties: **discrete fluctuation-dissipation balance (DFDB)**.

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- 2 Fluctuating Hydrodynamics
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- 4 Numerics**
- 5 Outlook

Numerical Scheme

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- The scheme ensures **strict conservation** of momentum and (almost exactly) enforces the no-slip condition at the end of the time step.
- Continuing work on temporal integrators that ensure the correct **equilibrium distribution** and **diffusive (Brownian) dynamics**.

Temporal Integrator (sketch)

- **Predict** particle position at midpoint:

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$$\begin{aligned} \rho \frac{\tilde{\mathbf{v}}^{n+1} - \mathbf{v}^n}{\Delta t} + \nabla \tilde{\pi} &= \frac{\eta}{2} \mathbf{L} (\tilde{\mathbf{v}}^{n+1} + \mathbf{v}^n) + \nabla \cdot \boldsymbol{\Sigma}^n + \mathbf{S}^{n+\frac{1}{2}} \mathbf{F}^{n+\frac{1}{2}} + \text{adv.}, \\ \nabla \cdot \tilde{\mathbf{v}}^{n+1} &= 0, \end{aligned}$$

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where we use the **Adams-Bashforth method** for the advective (kinetic) fluxes, and the discretization of the stochastic flux is described in Ref. [3],

$$\boldsymbol{\Sigma}^n = \left(\frac{2k_B T \eta}{\Delta V \Delta t} \right)^{1/2} \mathbf{W}^n,$$

where \mathbf{W}^n is a (symmetrized) collection of i.i.d. unit normal variates.

contd.

- Solve for inertial **velocity perturbation** from the particle $\Delta \mathbf{v}$ (too technical to present), and update:

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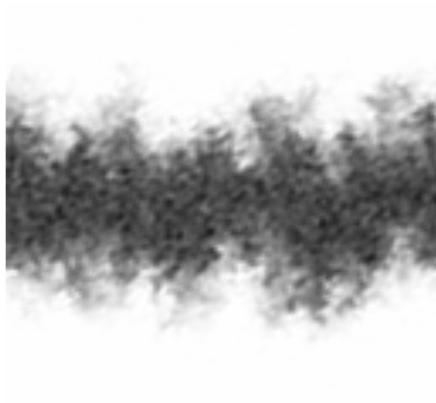
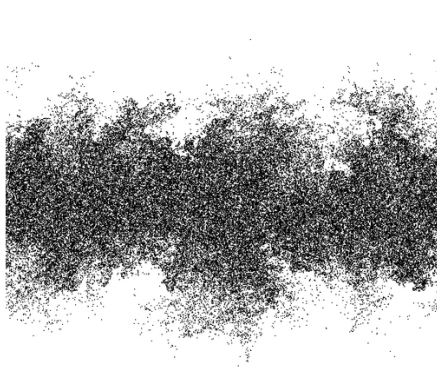
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Passively-Advected (Fluorescent) Tracers



Velocity Autocorrelation Function

- We investigate the **velocity autocorrelation function** (VACF) for the immersed particle

$$C(t) = \langle \mathbf{u}(t_0) \cdot \mathbf{u}(t_0 + t) \rangle$$

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- Hydrodynamic persistence (conservation) gives a **long-time power-law tail** $C(t) \sim (kT/m)(t/t_{\text{visc}})^{-3/2}$ not reproduced in Brownian dynamics.

Numerical VACF

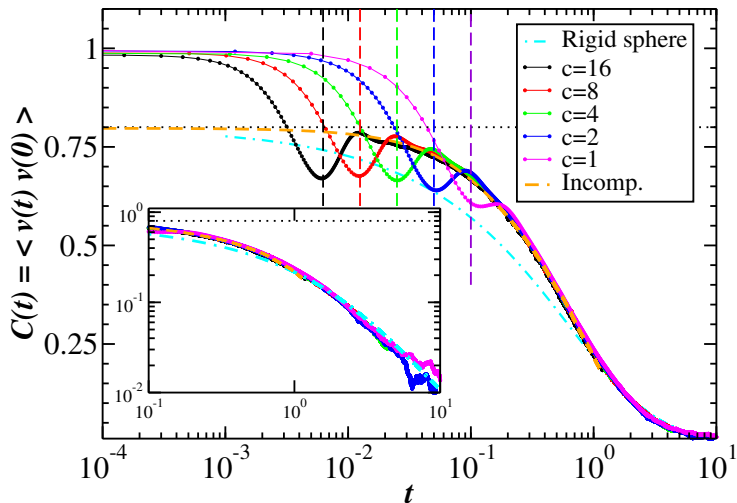


Figure: (F. Balboa) VACF for a blob with $m_e = m_f = \rho \Delta V$.

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- It is possible to include rotlet and stresslet terms, as done in the force coupling method [8] and Stokesian Dynamics in the deterministic setting.
- Proper inclusion of inertial terms and fluctuation-dissipation balance not studied carefully yet...

Immersed Rigid Bodies

- This approach can be extended to immersed rigid bodies (see work by Neelesh Patankar)

$$\rho (\partial_t \mathbf{v} + \mathbf{v} \cdot \nabla \mathbf{v}) = -\nabla \pi - \nabla \cdot \boldsymbol{\sigma} - \int_{\Omega} \mathbf{S}(\mathbf{q}) \boldsymbol{\lambda}(\mathbf{q}) d\mathbf{q} + \text{th. drift}$$

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- Even coarse-grained methods need to be accelerated due to **large separation of time scales** between advective and diffusive phenomena.
- One can take the **overdamped** (Brownian dynamics) **limit** but it would be much better to construct **many-scale temporal integrators** that are accurate even when they under-resolve the fast fluctuations.

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